

# **Нейтринные осцилляции: от измерений $\theta_{13}$ к иерархии масс и CP нарушению**

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**Сессия секции ЯФ ОФН РАН,  
ИФВЭ, Протвино, 5 ноября 2013**

# 2013 год: 100-летие со дня рождения Б.М. Понтецорво



## Mesonium and anti-mesonium

B. Pontecorvo

Sov.Phys.JETP 6 (1957) 429

Zh.Eksp.Teor.Fiz. 33 (1957) 549-551

## Inverse beta processes and nonconservation of lepton charge

B. Pontecorvo (Dubna, JINR)

Sov.Phys.JETP 7 (1958) 172-173,

Zh.Eksp.Teor.Fiz. 34 (1957) 247

## Neutrino Experiments and the Problem of Conservation of Leptonic Charge

B. Pontecorvo (Dubna, JINR)

Sov.Phys.JETP 26 (1968) 984-988,

Zh.Eksp.Teor.Fiz. 53 (1967) 1717-1725



# Outline

- neutrino mixing
- measurements of  $\theta_{13}$
- observation of  $\nu_e$  appearance
- near future: mixing parameters, MH and CP
- far future prospects
- summary

# $\nu$ oscillations and mixing

Standard Model: neutrinos are **massless** particles

3 families

atmospheric

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}$$

by June 2011

solar

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

link between atmospheric and solar

**$U$  parameterization:**

three mixing angles  $\theta_{12}$   $\theta_{23}$   $\theta_{13}$

CP violating phase  $\delta$

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

$$\Delta m_{12}^2 + \Delta m_{23}^2 + \Delta m_{31}^2 = 0$$

$$\theta_{23} \sim 45^\circ$$

$$|\Delta m_{23}^2| \approx |\Delta m_{31}^2| =$$

$$|\Delta m_{atm}^2| \approx 2.4 \times 10^{-3} \text{ eV}^2$$

$$\theta_{12} = (34 \pm 1)^\circ$$

$$\Delta m_{12}^2 = \Delta m_{sol}^2 \approx 7.5 \times 10^{-5} \text{ eV}^2$$

two independent  $\Delta m^2$

$$\sin^2 2\theta_{13} < 15^\circ \text{ at 90% CL}$$

??  $\theta_{13}$ , mass hierarchy,  $\delta$  ??



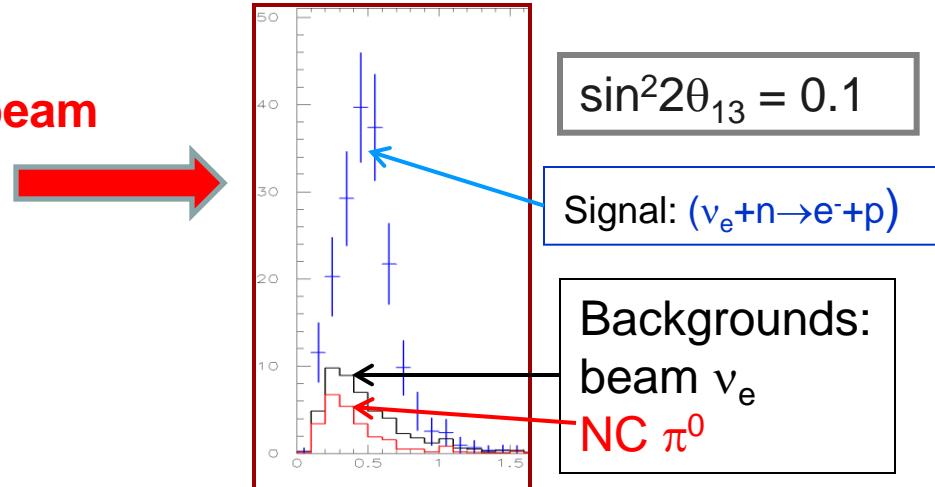
# Hunt for $\theta_{13}$

*Appearance:*

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 (\Delta m_{31}^2 L / 4E) + \text{CPV term} + \text{matter term} + \dots$$

Two methods

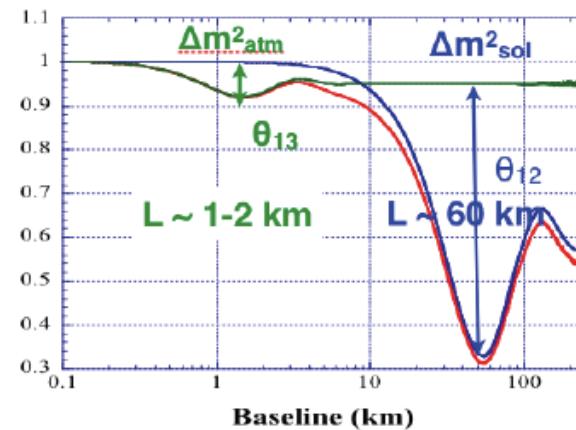
Search for electron neutrinos in beam  
of muon neutrinos  
at L/E tuned to the maximum  
of “atm” oscillations



*Disappearance:*

$$P(\nu_e \rightarrow \nu_e) \approx 1 - \sin^2 2\theta_{13} \sin^2 (\Delta m_{31}^2 L / 4E) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 (\Delta m_{21}^2 L / 4E)$$

Measurement of deficit of  
reactor antineutrinos  
at  $L \sim 1-2$  km





# Long-Baseline Neutrino Oscillation Experiment



SuperKamiokande

Toyama  
Kamioka Mine

~ 500 members  
59 institutions  
11 countries

JAPAN

Токио

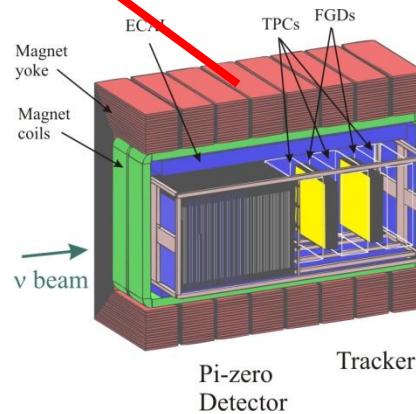
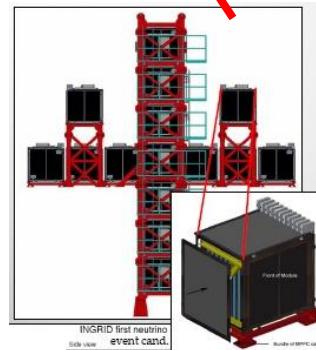
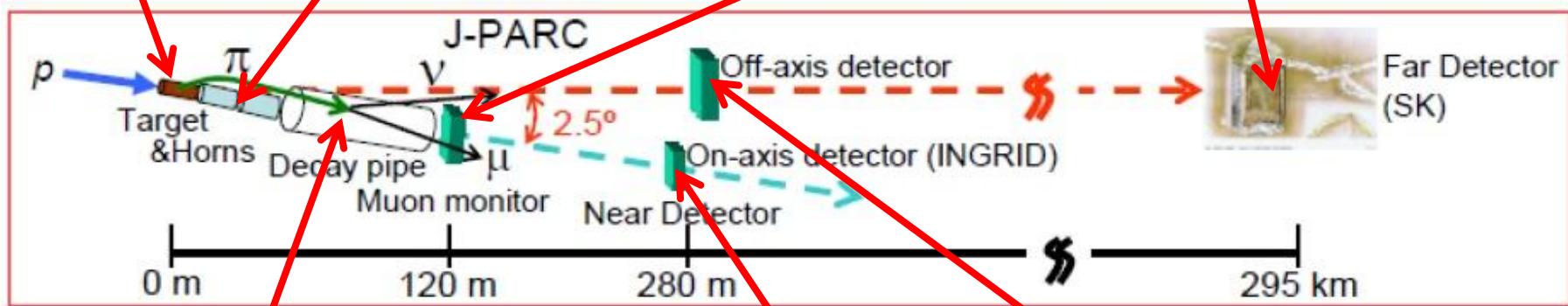
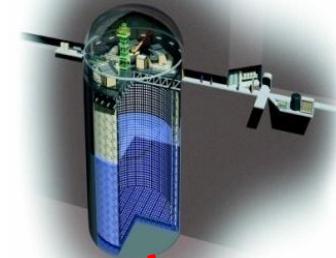
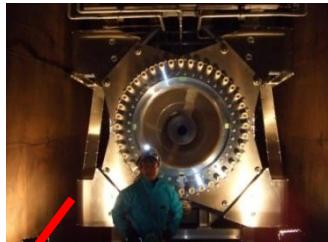


Tokai

Tokyo/Narita Airport



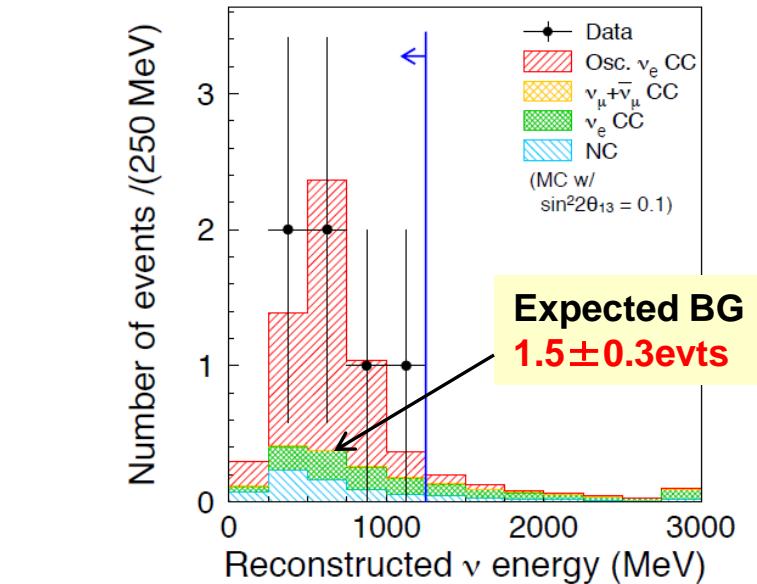
# T2K layout





# First T2K result

changed neutrino physics landscape



**1.43x10<sup>20</sup> POT**  
January 2010 –  
March 2011

In June 2011  
T2K published **first clear indication**  
of **electron neutrino appearance** ( $\theta_{13} \neq 0$ )

PRL 107, 041801 (2011) Selected for a Viewpoint in Physics

PHYSICAL REVIEW LETTERS

week ending  
22 JULY 2011

## Indication of Electron Neutrino Appearance from an Accelerator-Produced Off-Axis Muon Neutrino Beam

K. Abe,<sup>49</sup> N. Abgrall,<sup>16</sup> Y. Ajima,<sup>18,†</sup> H. Aihara,<sup>48</sup> J. B. Albert,<sup>13</sup> C. Andreopoulos,<sup>47</sup> B. Andrieu,<sup>37</sup> S. Aoki,<sup>27</sup> O. Araoka,<sup>18,†</sup> J. Argyriades,<sup>16</sup> A. Ariga,<sup>3</sup> T. Ariga,<sup>3</sup> S. Assylbekov,<sup>11</sup> D. Autiero,<sup>32</sup> A. Badertscher,<sup>15</sup> M. Barbi,<sup>40</sup> G. J. Barker,<sup>56</sup> G. Barr,<sup>36</sup> M. Bass,<sup>11</sup> F. Bay,<sup>3</sup> S. Bentham,<sup>29</sup> V. Berardi,<sup>22</sup> B. E. Berger,<sup>11</sup> I. Bertram,<sup>29</sup> M. Besnier,<sup>14</sup> J. Beucher,<sup>8</sup> D. Beznosko,<sup>34</sup> S. Bhadra,<sup>59</sup> F. d. M. M. Blaszczyk,<sup>8</sup> A. Blondel,<sup>16</sup> C. Bojeckho,<sup>53</sup> J. Bouchez,<sup>8,\*</sup> S. B. Boyd,<sup>56</sup> A. Bravar,<sup>16</sup> C. Bronner,<sup>14</sup> D. G. Brook-Roberge,<sup>5</sup> N. Buchanan,<sup>11</sup> H. Budd,<sup>41</sup> D. Calvet,<sup>8</sup> S. L. Cartwright,<sup>44</sup> A. Carver,<sup>56</sup> R. Castillo,<sup>19</sup> M. G. Catanese,<sup>22</sup> A. Cazes,<sup>32</sup> A. Cervera,<sup>20</sup> C. Chavez,<sup>30</sup> S. Choi,<sup>43</sup> G. Christodoulou,<sup>30</sup> J. Coleman,<sup>30</sup>

The T2K experiment observes indications of  $\nu_\mu \rightarrow \nu_e$  appearance in data accumulated with  $1.43 \times 10^{20}$  protons on target. Six events pass all selection criteria at the far detector. In a three-flavor neutrino oscillation scenario with  $|\Delta m_{23}^2| = 2.4 \times 10^{-3}$  eV<sup>2</sup>,  $\sin^2 2\theta_{23} = 1$  and  $\sin^2 2\theta_{13} = 0$ , the expected number of such events is  $1.5 \pm 0.3$  (syst). Under this hypothesis, the probability to observe six or more candidate events is  $7 \times 10^{-3}$ , equivalent to  $2.5\sigma$  significance. At 90% C.L., the data are consistent with  $0.03(0.04) < \sin^2 2\theta_{13} < 0.28(0.34)$  for  $\delta_{CP} = 0$  and a normal (inverted) hierarchy.

DOI: 10.1103/PhysRevLett.107.041801

PACS numbers: 14.60.Pq, 13.15.+g, 25.30.Pt, 95.55.Vj

Then,  
1 - Confirmation from MINOS  
2 - Precise measurements by  
Double Chooz  
Daya Bay  
RENO

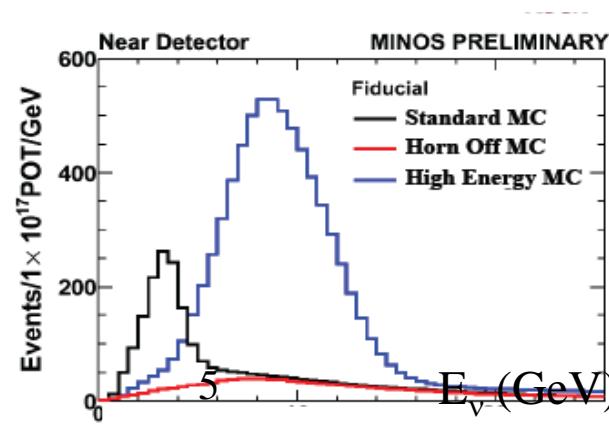


## Far detector



# MINOS

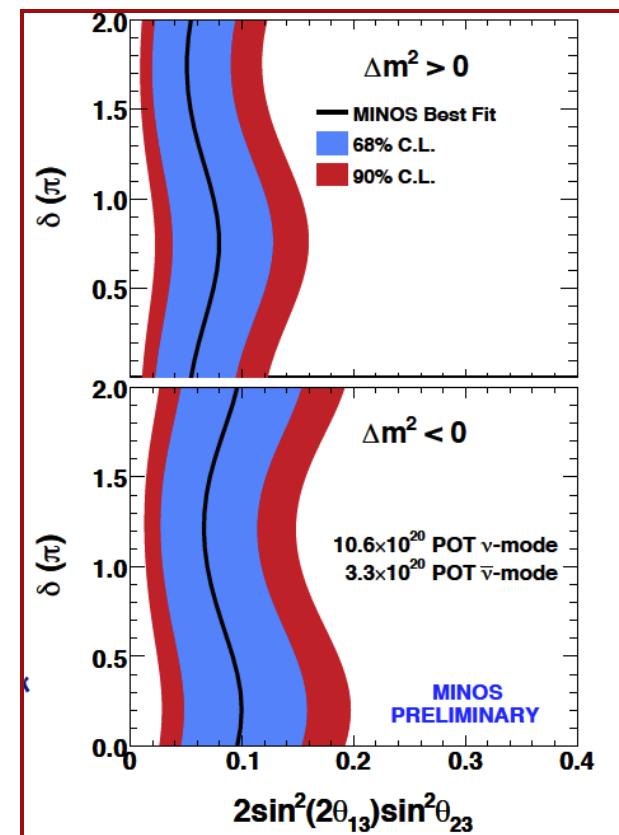
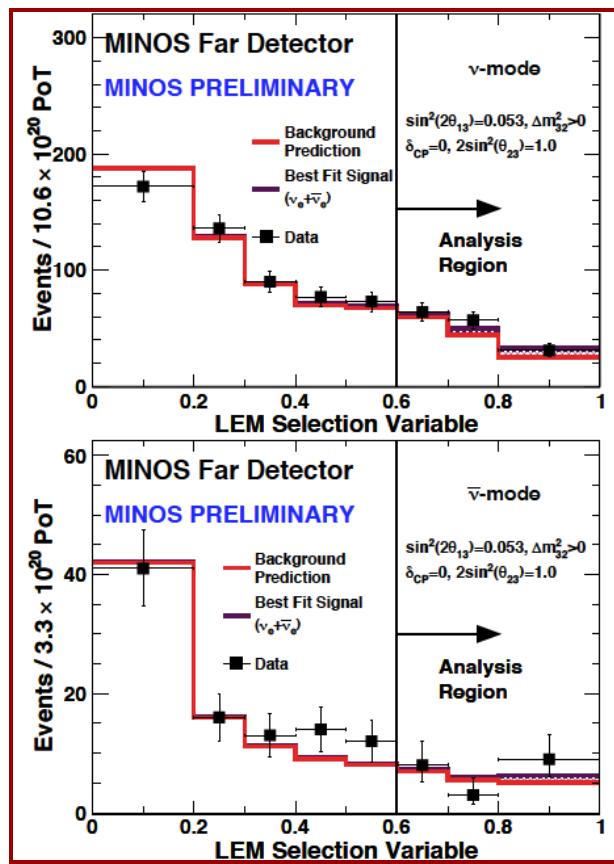
- ~3GeV  $\nu_\mu$  beam from FNAL 120GeV MI
  - ~350kW operation achieved
- (magnetized)Iron-scintillator tracker at 735km (5.4kt) and near (980t)
- Main physics goals
  - (anti-) $\nu_\mu$  disappearance
  - $\nu_e$  appearance
- **Integral luminosity**
  - $10.7 \times 10^{20}$  POT for  $\nu_\mu$
  - $3.4 \times 10^{20}$  POT for anti- $\nu_\mu$





# MINOS: $\theta_{13}$

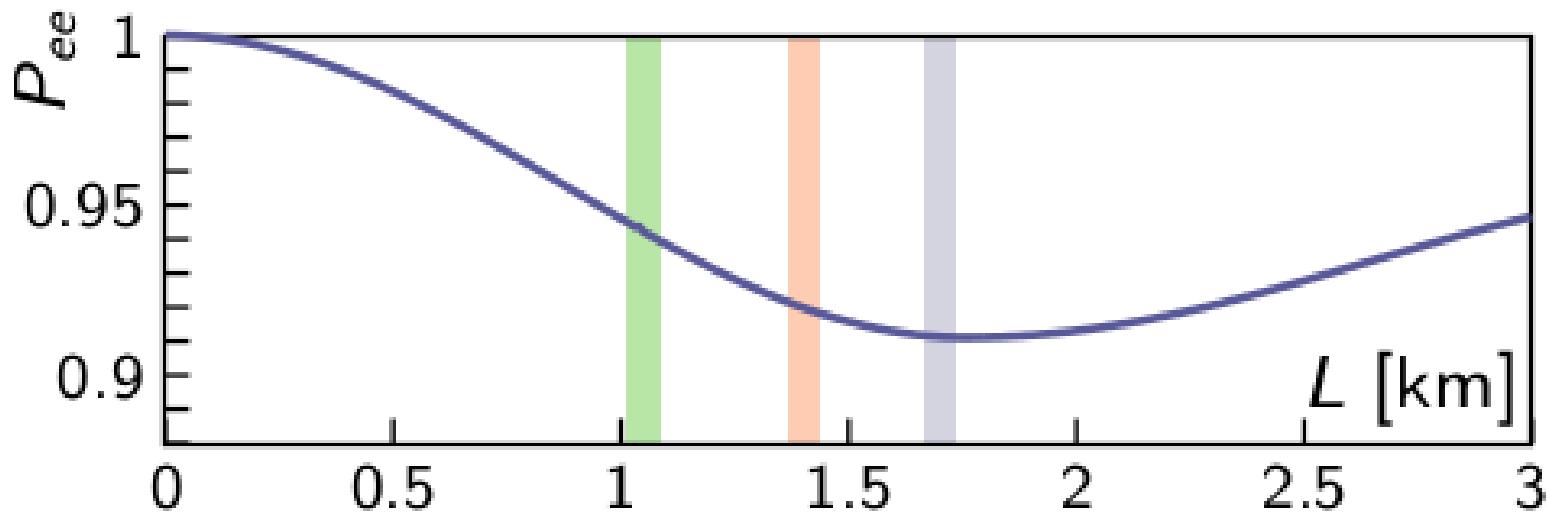
- Neutrino beam
  - Expect: 128.6(+32.5) events
  - Observe: 152 Events
- Antineutrino beam
  - Expect 17.5(+3.7) events
  - Observe 20 events



disfavour  $\theta_{13}=0$  at 96% CL  
For normal mass hierarchy and  $\delta=0$

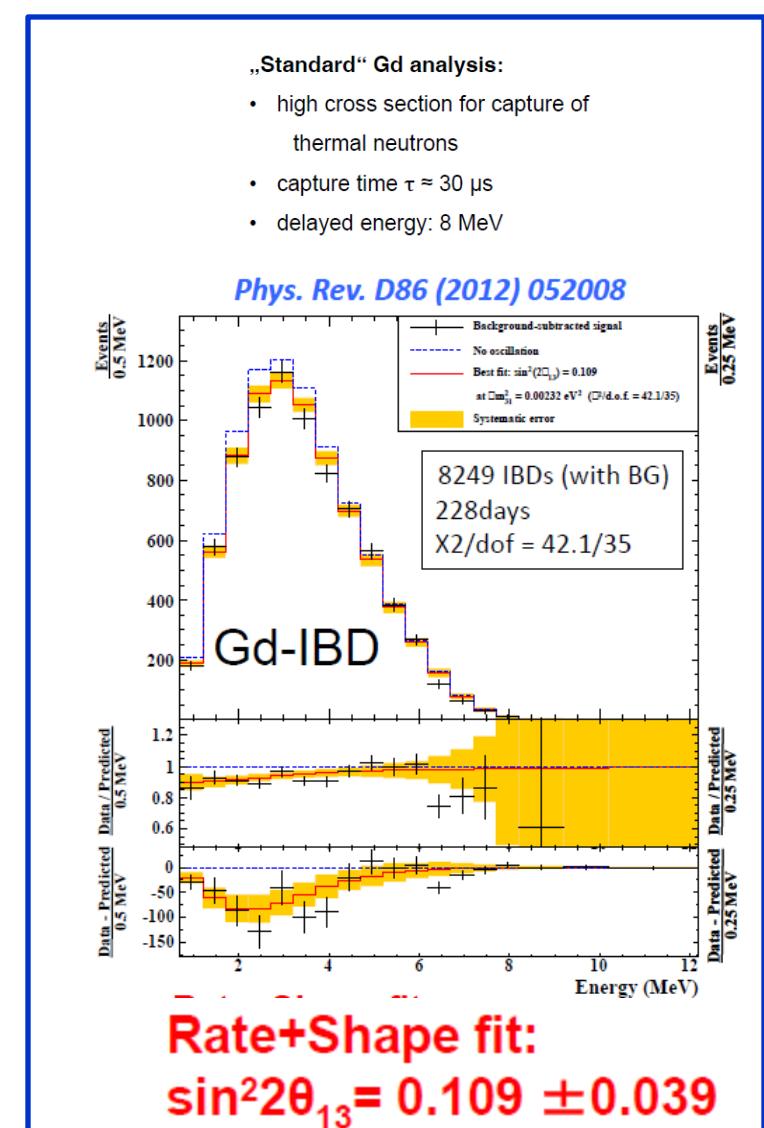
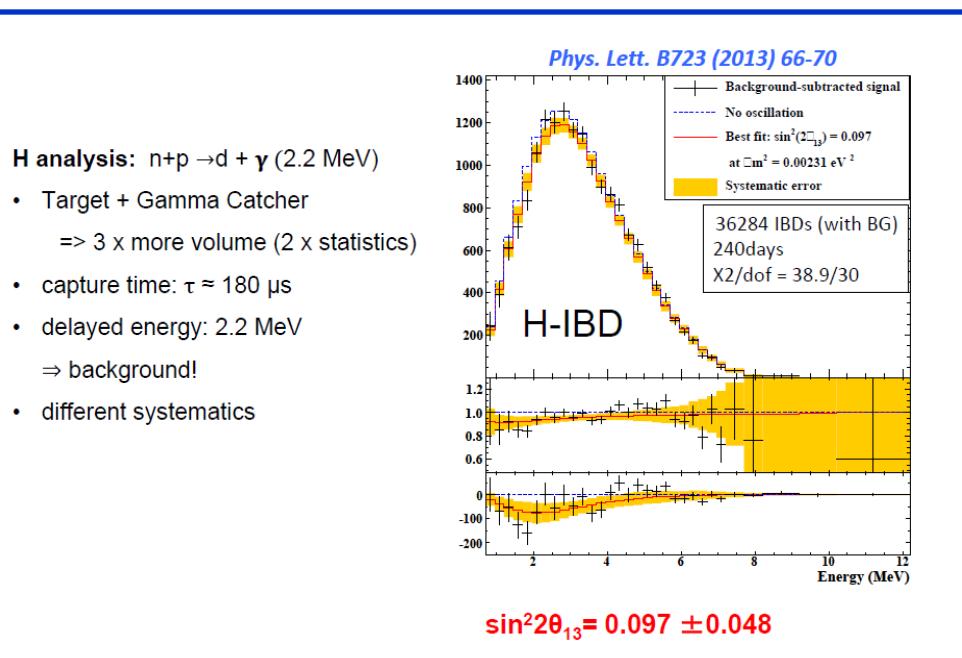
# $\theta_{13}$ from reactor experiments

## 3 reactor experiments: Double Chooz, Daya Bay, RENO



	Reactor [GW <sub>th</sub> ]	Target [tons]	Depth [m.w.e]
<b>Double Chooz</b>	8.6	16 (2 × 8)	300, 120 (far, near)
<b>RENO</b>	16.5	32 (2 × 16)	450, 120
<b>Daya Bay</b>	<b>17.4</b>	<b>160 (8 × 20)</b>	<b>860, 250</b>
Large Signal			Low Background

# Double Chooz



**Combined:**  
 $\sin^2 2\theta_{13} = 0.109 \pm 0.035$



# Daya Bay

Reactor experiment in China



- 17.4 GW<sub>th</sub> power
- 8 operating detectors
- 160 t total target mass

$3.5 \times 10^{21}$  neutrinos per second



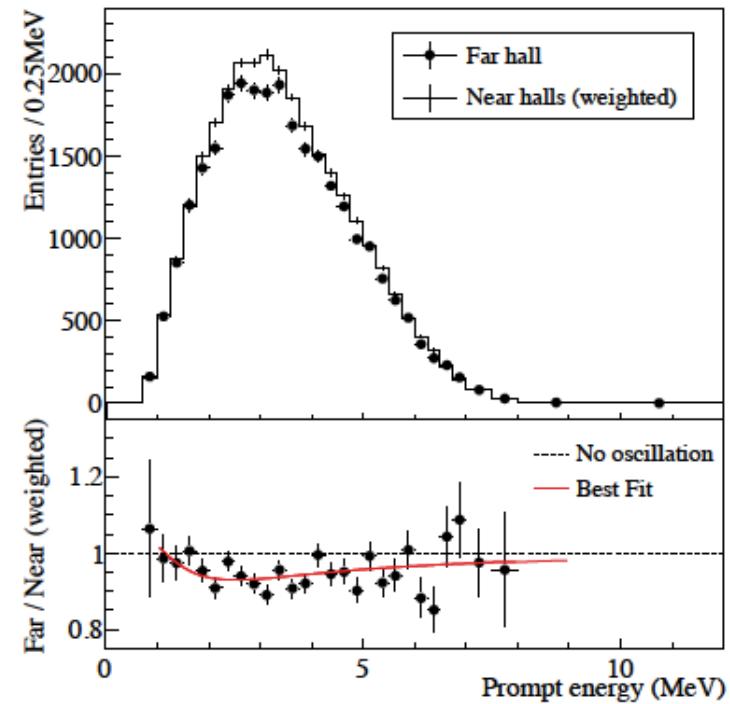
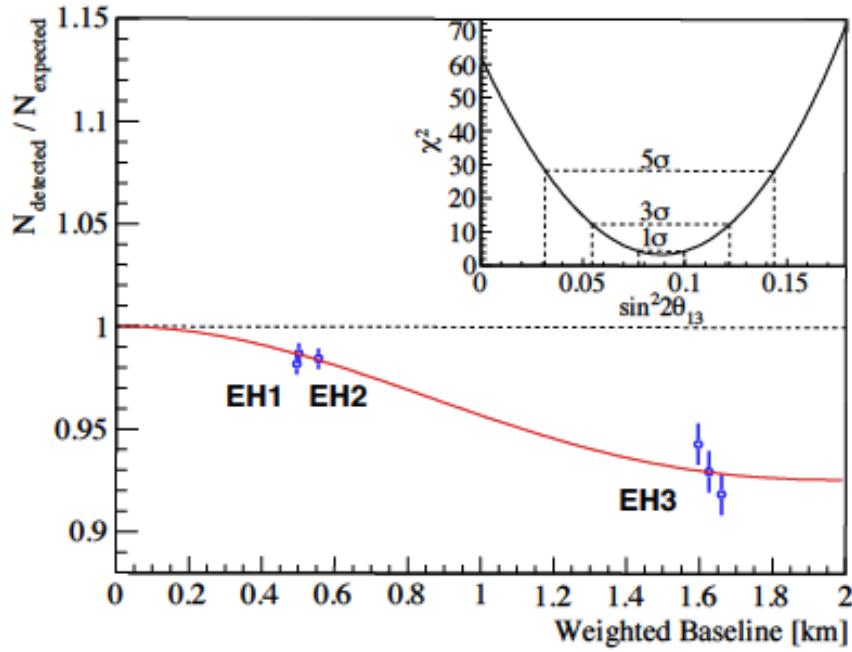
	Overburden	$R_\mu$	$E_\mu$	D1,2	L1,2	L3,4
EH1	250	1.27	57	364	857	1307
EH2	265	0.95	58	1348	480	528
EH3	860	0.056	137	1912	1540	1548

TABLE I. Overburden (m.w.e), muon rate  $R_\mu$  (Hz/m<sup>2</sup>), and average muon energy  $E_\mu$  (GeV) of the three EHs, and the distances (m) to the reactor pairs.



# Daya Bay result (I)

March 2012, after 55 days of data taking



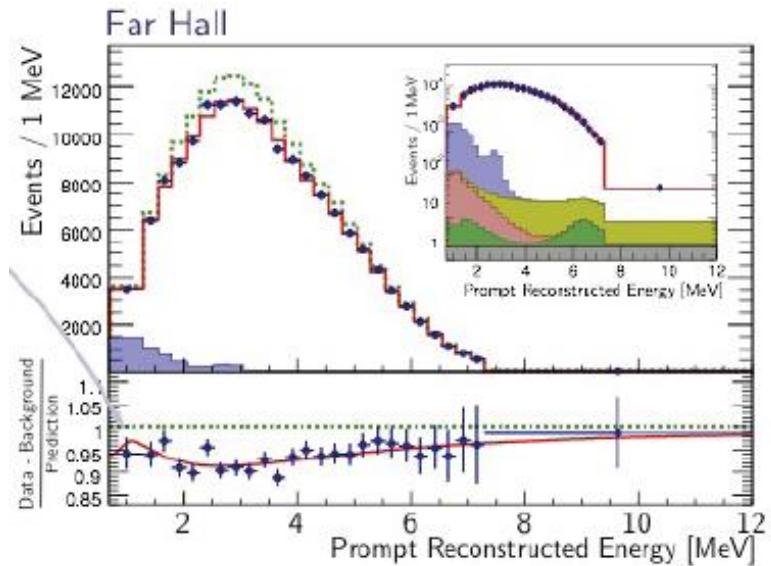
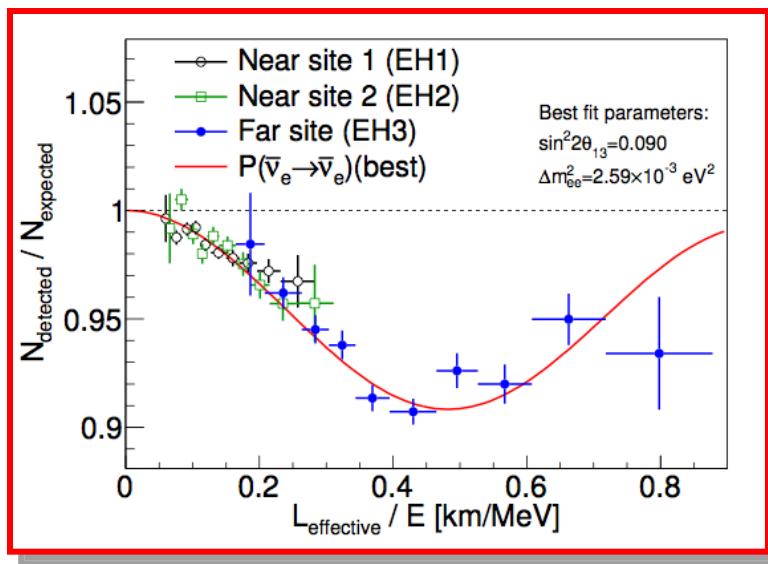
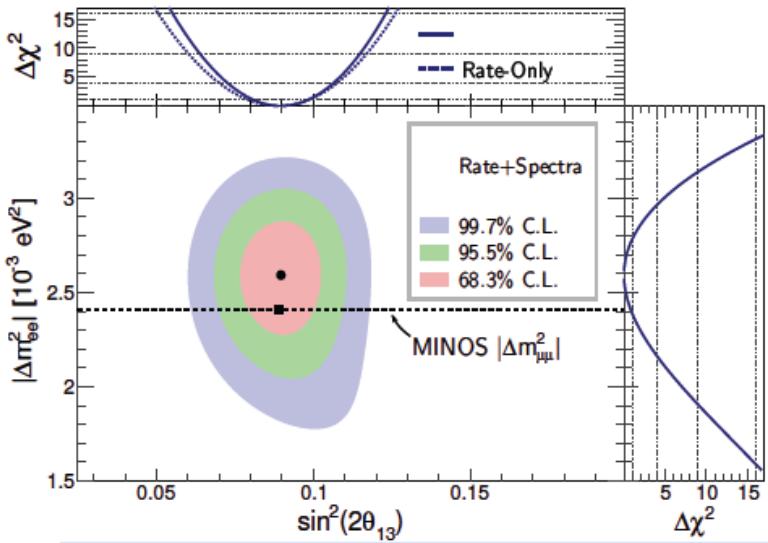
$$\sin^2 2\theta_{13} = 0.089 \pm 0.010 \text{ (stat)} \pm 0.005 \text{ (syst)}$$



# Daya Bay result (II)

Rate and spectral analysis

217 days of data taking

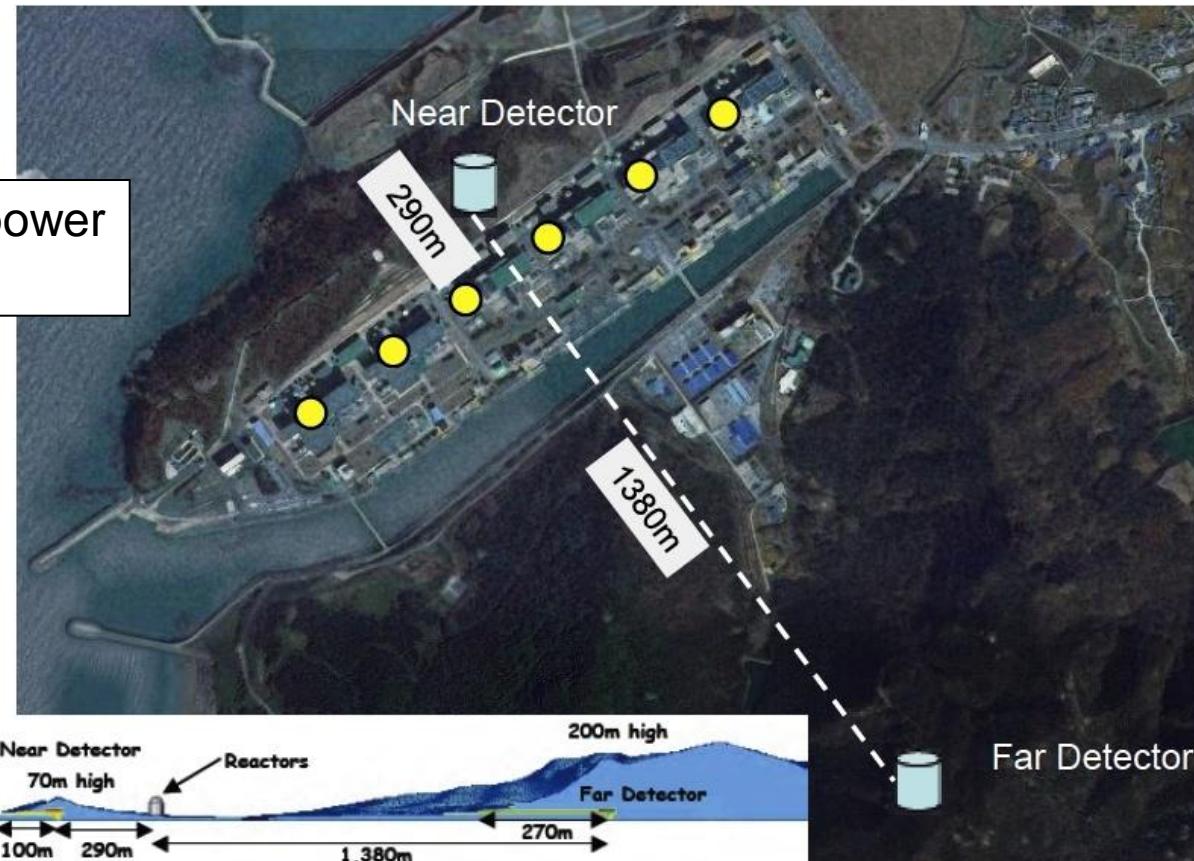


$$\sin^2 2\theta_{13} = 0.090^{+0.008}_{-0.009}$$

$$|\Delta m_{ee}^2| = 2.59^{+0.19}_{-0.20} \times 10^{-3} \text{ eV}^2$$

# RENO

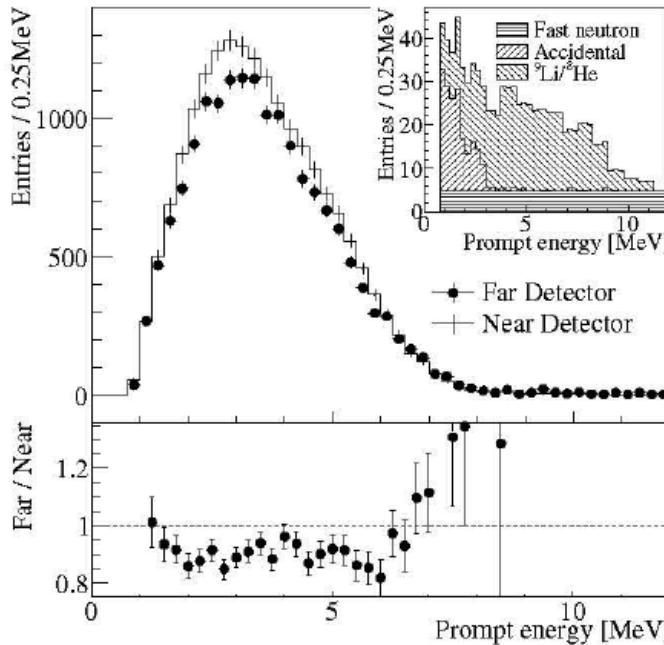
Reactor experiment in Korea



- Target : 16.5 ton Gd-LS, R=1.4m, H=3.2m
- Gamma Catcher : 30 ton LS, R=2.0m, H=4.4m
- Buffer : 65 ton mineral oil, R=2.7m, H=5.8m
- Veto : 350 ton water, R=4.2m, H=8.8m

Data taking since 1<sup>st</sup> August 2011  
First result published in May 2012

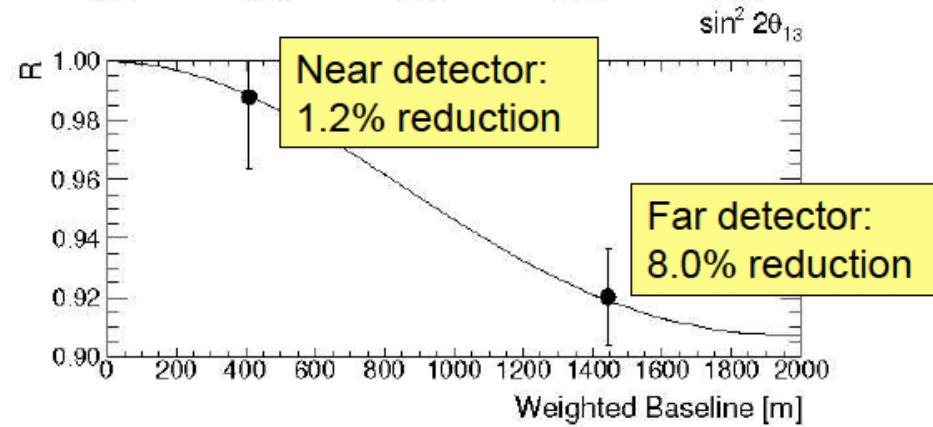
# RENO result (I)



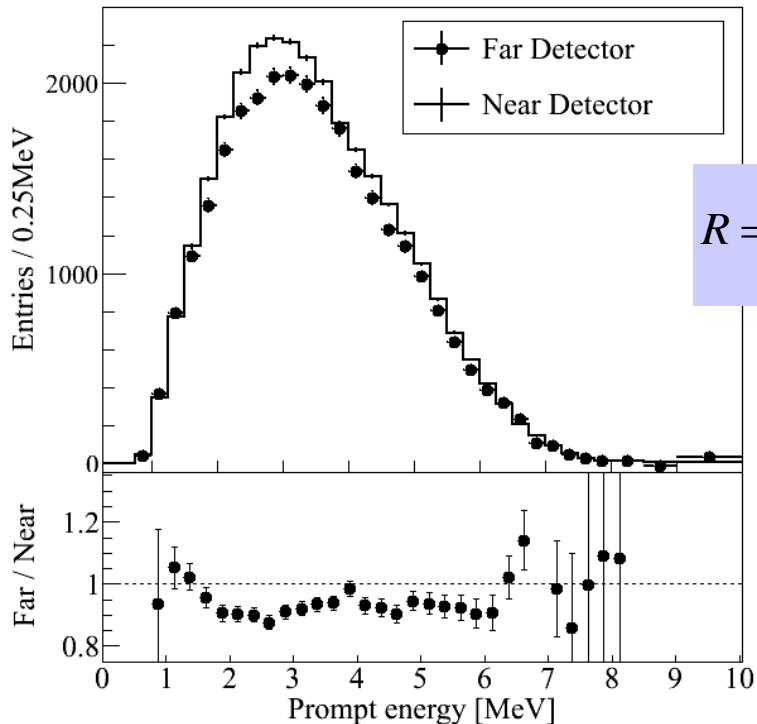
$$R = \frac{\Phi_{\text{observed}}^{\text{Far}}}{\Phi_{\text{expected}}^{\text{Far}}} = 0.920 \pm 0.009(\text{stat}) \pm 0.014(\text{syst})$$

$$\sin^2 2\theta_{13} = 0.113 \pm 0.013(\text{stat.}) \pm 0.019(\text{syst.})$$

significance  $4.9\sigma$



# RENO result (II)



S.-B.Kim  
Erice/Sicily, Sep. 2013

$$R = \frac{\Phi_{\text{observed}}^{\text{Far}}}{\Phi_{\text{expected}}^{\text{Far}}} = 0.929 \pm 0.006(\text{stat}) \pm 0.007(\text{syst})$$

- A clear deficit in rate (~ 7 % reduction)
- Consistent with neutrino oscillation in the spectral distortion

- a new result on mixing angle  $\theta_{13}$ .

$$\sin^2 2\theta_{13} = 0.100 \pm 0.010(\text{stat}) \pm 0.012(\text{syst}) \quad (\text{402 days data taking})$$

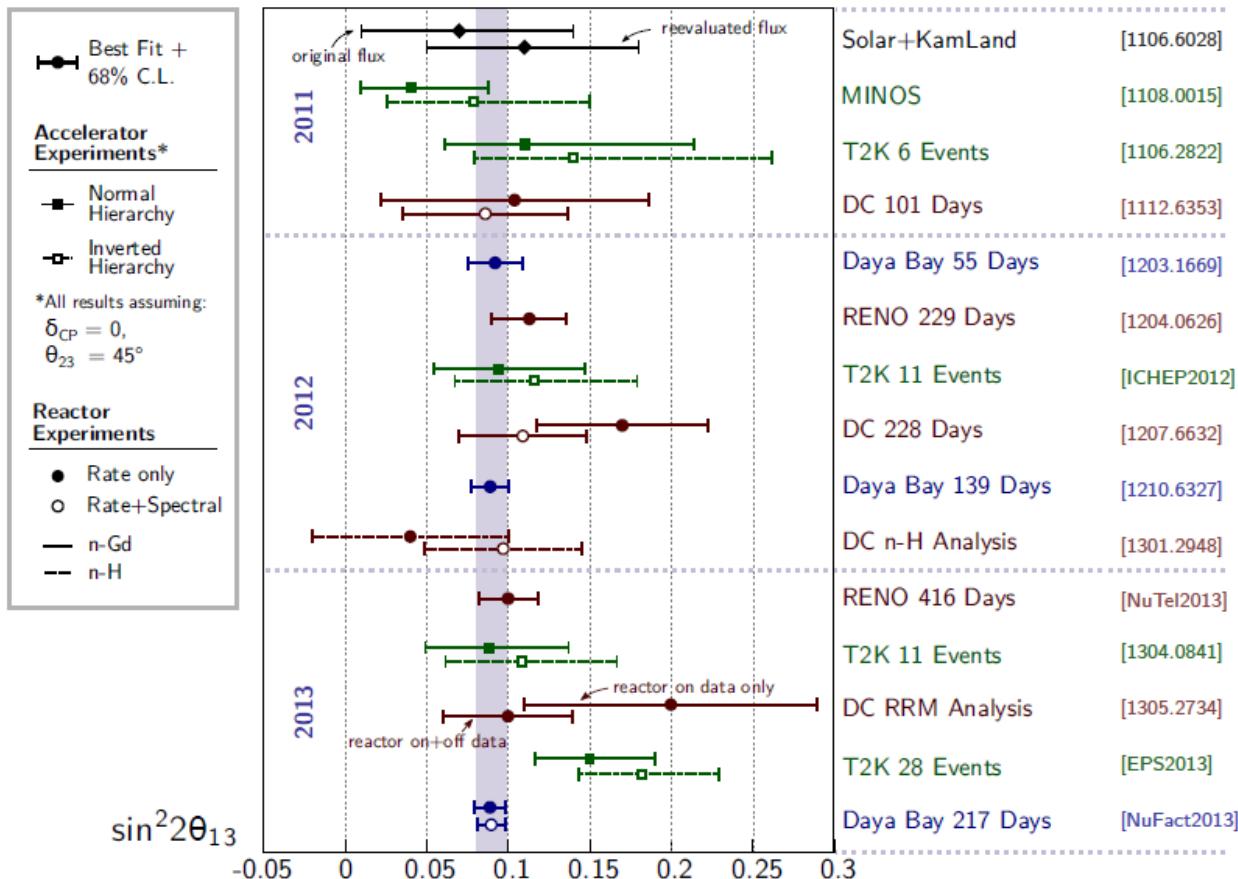
(systematic error will be further reduced)



# History of $\theta_{13}$ measurements

Only 2 years of measurements!

S. Jetter, NuFact2013



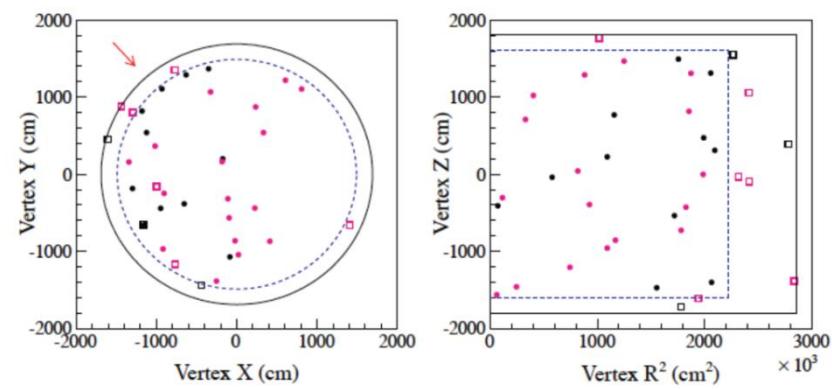
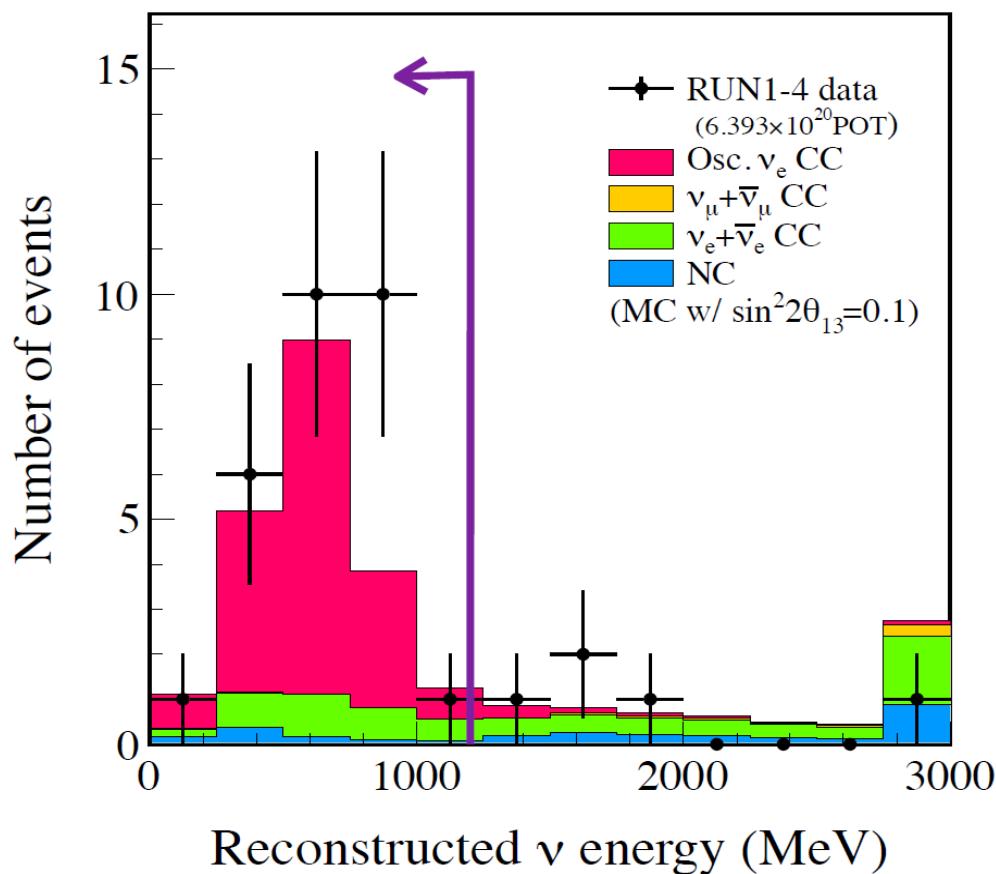
$$\theta_{13} \approx 90^\circ$$



# T2K: observation of $\nu_e$ appearance

Summer 2013: integral luminosity  $6.39 \times 10^{20}$  pot

28  $\nu_e$  candidates

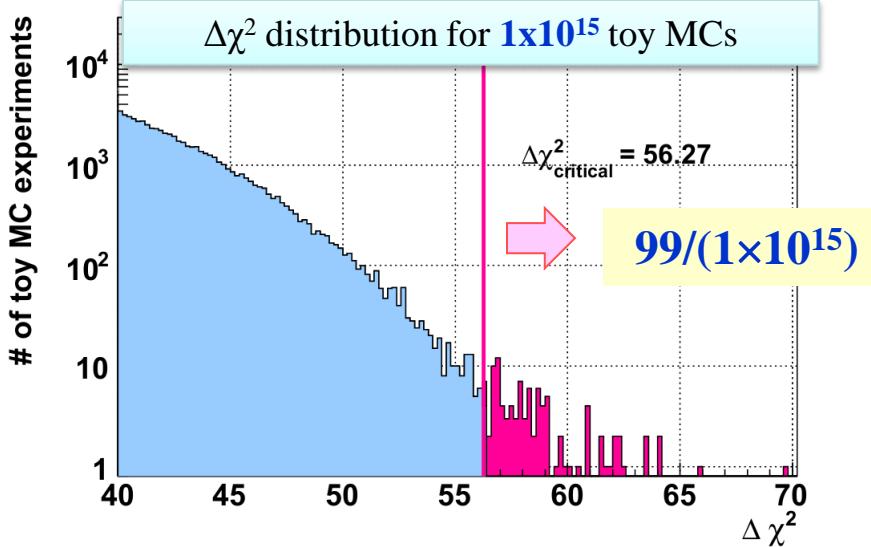
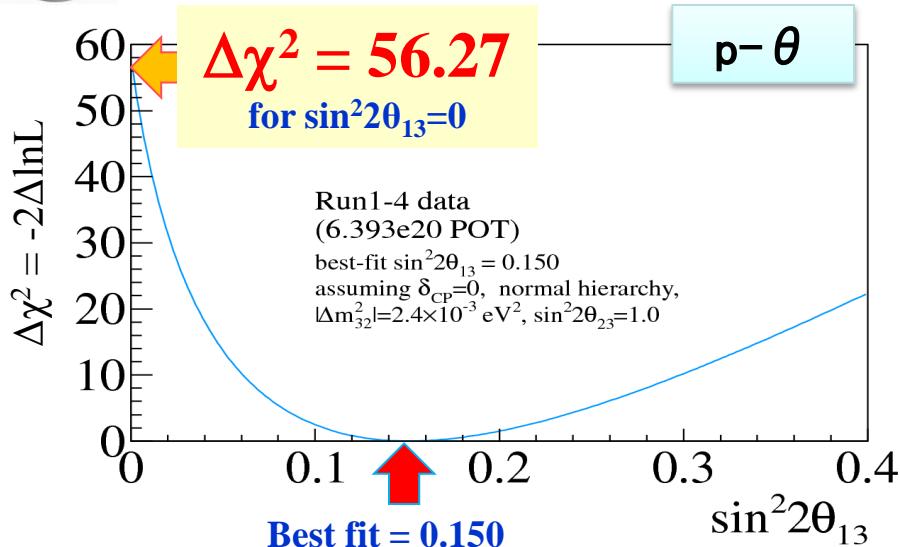


Expected background  
( $\sin^2 2\theta_{13} = 0.0$ )  
 $4.64 \pm 0.52$  events

# Significance



significance is calculated as  $\sqrt{\Delta\chi^2}$



28  $\nu_e$  events detected  
expected background  
 $4.64 \pm 0.52$  events

$$\sqrt{-2\Delta\ln L} = \sqrt{56.27} = 7.5\sigma$$

p-value is calculated as follows:

1. Generate  $1e15$  toy experiments with  $\sin^2 2\theta_{13}=0.0$ .
2. Fit each toy experiment extract  $-2\Delta\ln L$  ( $=\Delta\chi^2$ ).
3. p-value is the fraction of toy experiments above  $\Delta\chi^2_{data}$

p-value =  $9.9 \times 10^{-14}$

Discovery of  $\nu_e$  appearance



# Result: $\theta_{13}$ vs $\delta$

allowed region of  $\sin^2 2\theta_{13}$   
for each value of  $\delta_{CP}$

Best fit for  $\delta_{CP}=0$

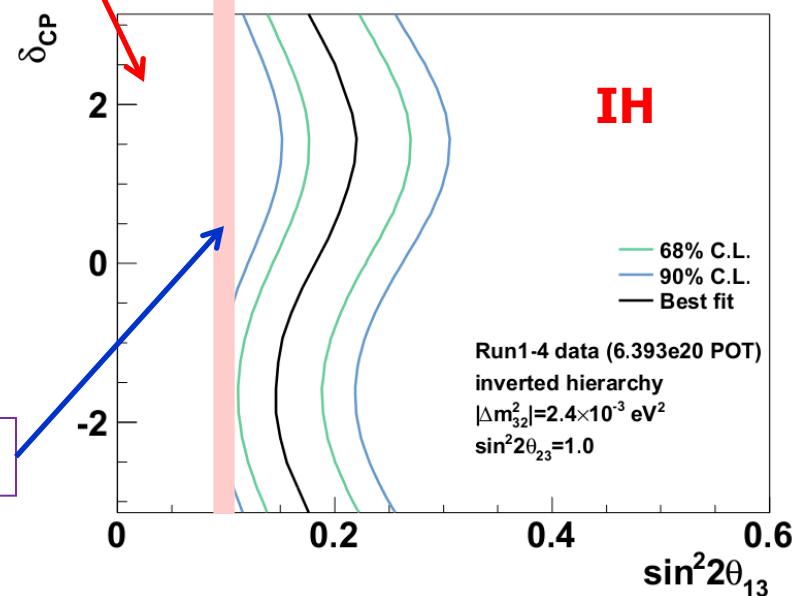
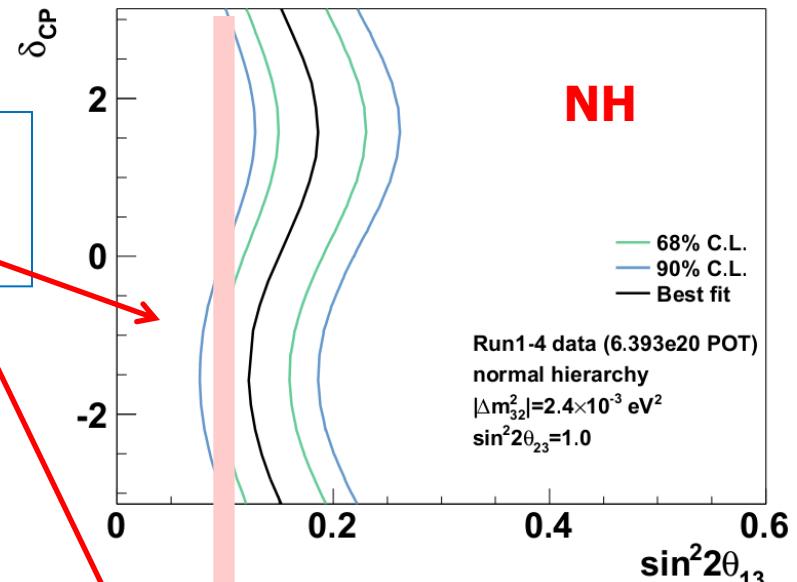
NH

$$\sin^2 2\theta_{13} = 0.150^{+0.039}_{-0.034}$$

IH

$$\sin^2 2\theta_{13} = 0.182^{+0.046}_{-0.040}$$

PDG(2012):  $\sin^2 2\theta_{13} : 0.098 \pm 0.013$





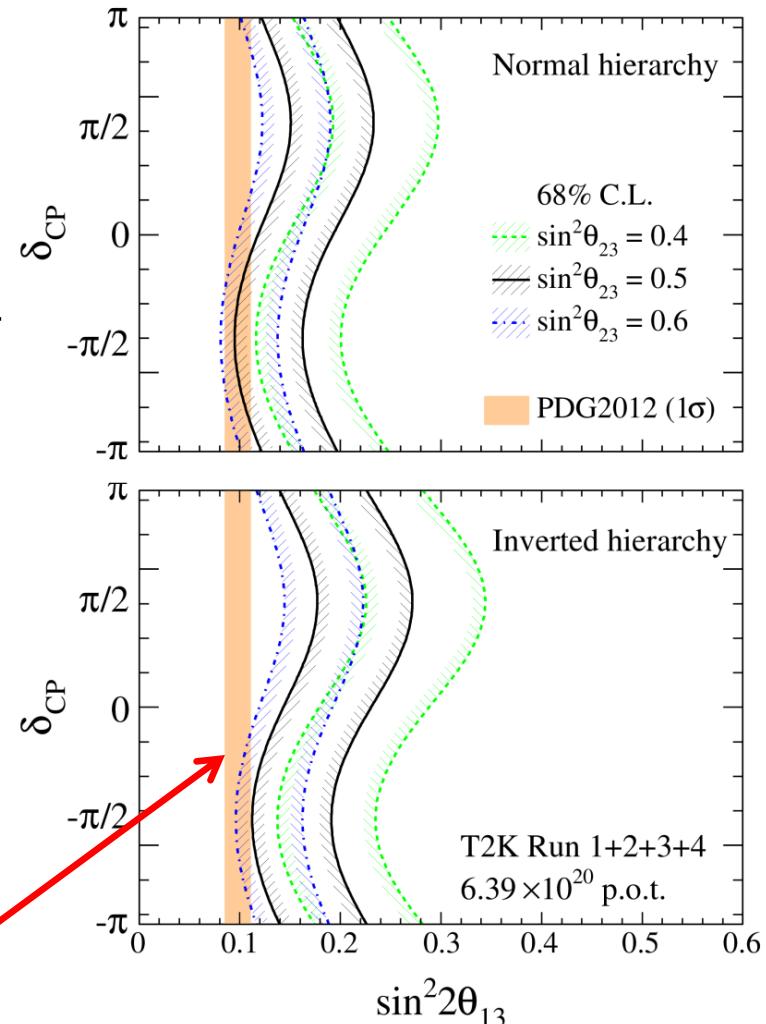
# $\theta_{23}$ issue

$\theta_{23}$  uncertainties  
dominate in  
 $\delta - \sin^2 2\theta_{13}$  plot

$$P_{\nu_\mu \rightarrow \nu_e} \approx [\sin^2 \theta_{23}] \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{32}^2 L}{4E_\nu}$$

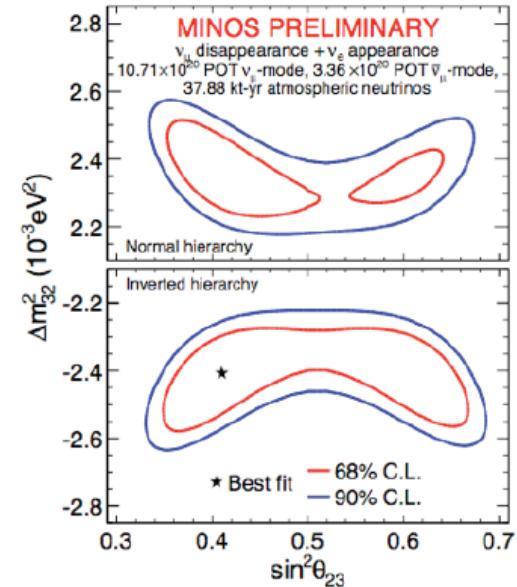
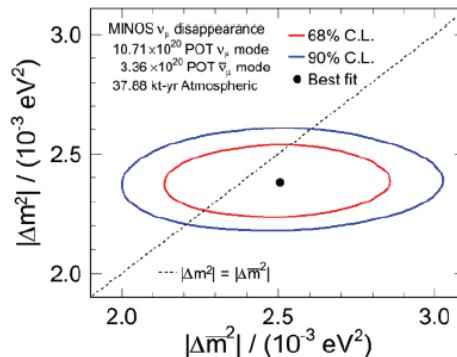
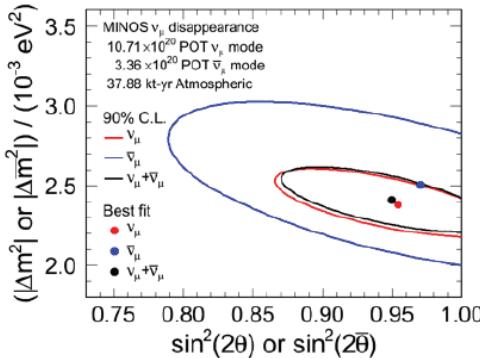
- Oscillation probability is dependent on  $\sin^2 \theta_{23}$  (octant)
  - PDG2012:  $\sin^2(2\theta_{23}) > 0.95$ 
    - $\sin^2 \theta_{23} = 0.50 \pm 0.11$
    - $\theta_{23} = 45 \pm 6.5^\circ$
  - Reduction of  $\sin^2 \theta_{23}$  error is critical for further improvements

PDG(2012)





# MINOS



## $\bar{\nu}$ oscillation parameters

$$\sin^2(2\bar{\theta}) = 0.97^{+0.03}_{-0.08}$$

$$\Delta \bar{m}^2 = 2.50^{+0.23}_{-0.25} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\bar{\theta}) > 0.83 \text{ (90\% C.L.)}$$

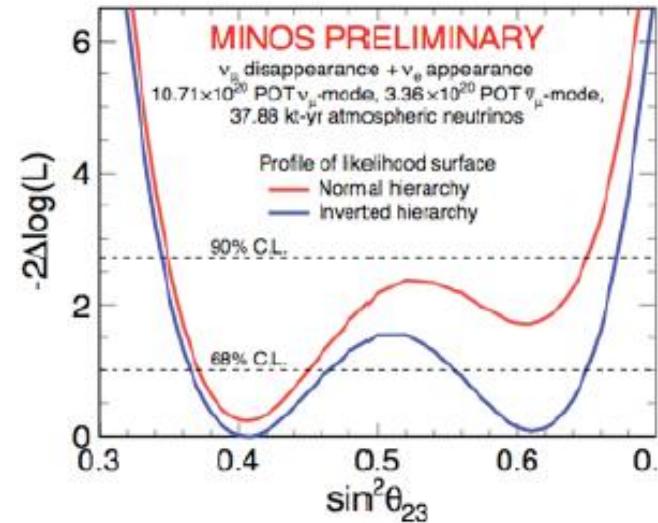
## $\nu$ oscillation parameters

$$\sin^2(2\theta) = 0.95^{+0.035}_{-0.036}$$

$$|\Delta m^2| = 2.41^{+0.09}_{-0.10} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta) > 0.89 \text{ (90\% C.L.)}$$

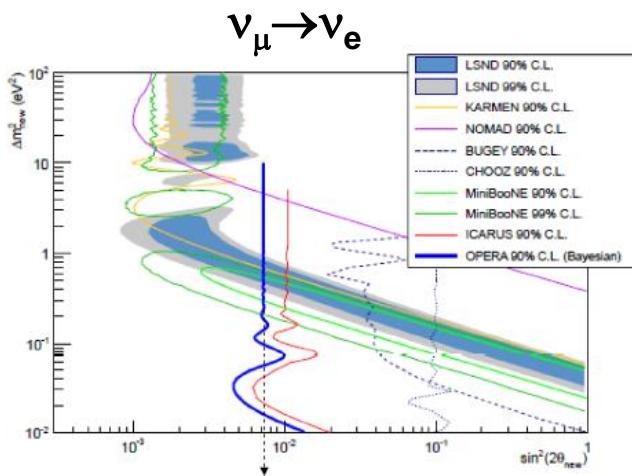
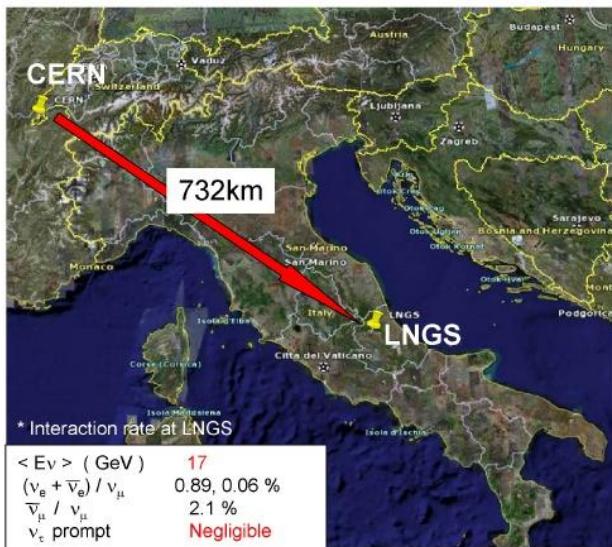
**MINOS Preferences:**  
**Low octant,  $\theta_{23} < 45^\circ$**   
**Non-max mixing**  
**Inverted Mass hierarchy**



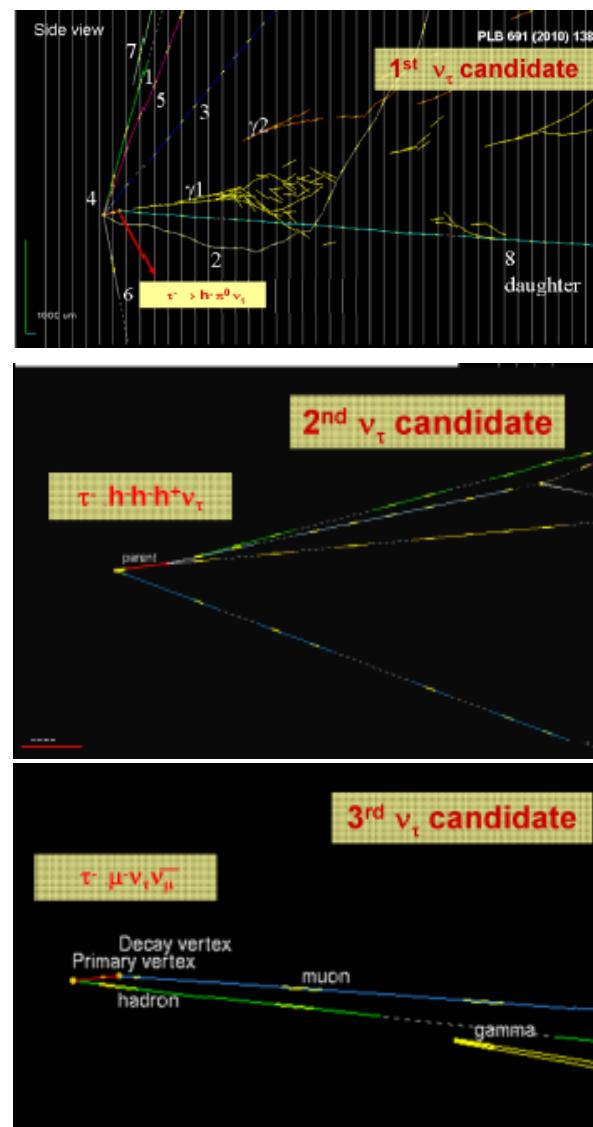


# OPERA

Neutrino beam from CERN to Gran Sasso

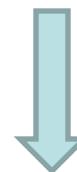


$\nu_\mu \rightarrow \nu_\tau$



Data  
2008-2009

Expected bkg  
0.226  $\nu_\tau$  events



3  $\nu_\tau$  events:  
significance

**3.2σ**

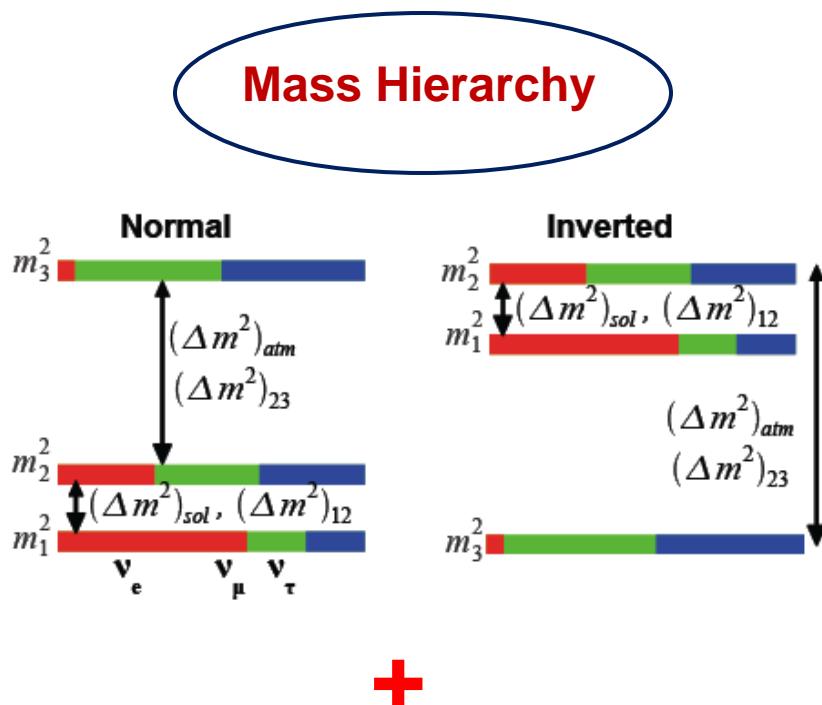
p-value of  
background  
 **$7.3 \times 10^{-3}$**



# Next targets ?

**2011-2013  
Great milestones**

- $\theta_{13} = (9 \pm 0.6)^0$
- $\nu_e$  appearance observed at  $7.5\sigma$



+  
precision measurement of  
oscillation parameters

$\theta_{23}$ :  $\geq 45^\circ$  or  $\leq 45^\circ$

$$J_{CP} = \text{Im}(U_{e1} U_{\mu 2} U_{e2}^* U_{\mu 1}^*) = \text{Im}(U_{e2} U_{\mu 3} U_{e3}^* U_{\mu 2}^*) \\ = \cos\theta_{12} \sin\theta_{12} \cos^2\theta_{13} \sin\theta_{13} \cos\theta_{23} \sin\theta_{23} \sin\delta$$

all mixing angles  $\neq 0 \rightarrow J_{CP} \neq 0$  if  $\delta \neq 0$

CKM

$$\begin{pmatrix} 1 & 0.2 & 0.004 \\ 0.2 & 1 & 0.04 \\ 0.008 & 0.04 & 1 \end{pmatrix}$$

PMNS

$$\begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

Quark sector  $J_{CP} \approx 3 \times 10^{-5}$

Lepton sector  $J_{CP} \sim 0.02 \times \sin\delta$



# $\nu_\mu \rightarrow \nu_e$ in matter

Physics reach oscillation mode for accelerator LBL experiments is  $\nu_\mu \rightarrow \nu_e$

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4c_{13}^2 s_{13}^2 s_{23}^2 \sin^2 \frac{\Delta m_{13}^2 L}{4E_\nu} \times \left[ 1 + \frac{2a}{\Delta m_{13}^2} (1 - 2s_{13}^2) \right] \xrightarrow{\theta_{13}} \\
 & + 8c_{13}^2 s_{12} s_{13} s_{23} (c_{12} c_{23} \cos \delta - s_{12} s_{13} s_{23}) \cos \frac{\Delta m_{23}^2 L}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} \sin \frac{\Delta m_{12}^2 L}{4E_\nu} \xrightarrow{\text{CP-even}} \\
 & - 8c_{13}^2 c_{12} c_{23} s_{12} s_{13} s_{23} \sin \delta \sin \frac{\Delta m_{23}^2 L}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} \sin \frac{\Delta m_{12}^2 L}{4E_\nu} \xrightarrow{\text{CP-odd}} \\
 & + 4s_{12}^2 c_{13}^2 (c_{13}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{23} s_{13} \cos \delta) \sin^2 \frac{\Delta m_{12}^2 L}{4E_\nu} \xrightarrow{\text{Solar}} \\
 & - 8c_{13}^2 s_{13}^2 s_{23}^2 \cos \frac{\Delta m_{23}^2 L}{4E_\nu} \frac{aL}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} (1 - 2s_{13}^2), \xrightarrow{(30)} \xrightarrow{\text{Matter}}
 \end{aligned}$$

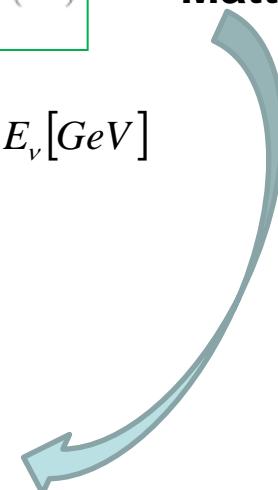
$$s_{ij} = \sin \theta_{ij} \quad c_{ij} = \cos \theta_{ij} \quad a [eV^2] = 2\sqrt{2}G_F n_e E_\nu = 7.6 \times 10^{-5} \rho \left[ \frac{g}{cm^3} \right] E_\nu [GeV]$$

$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$



$a \rightarrow -a \quad \delta \rightarrow -\delta$

change sign for NH  $\rightarrow$  IH

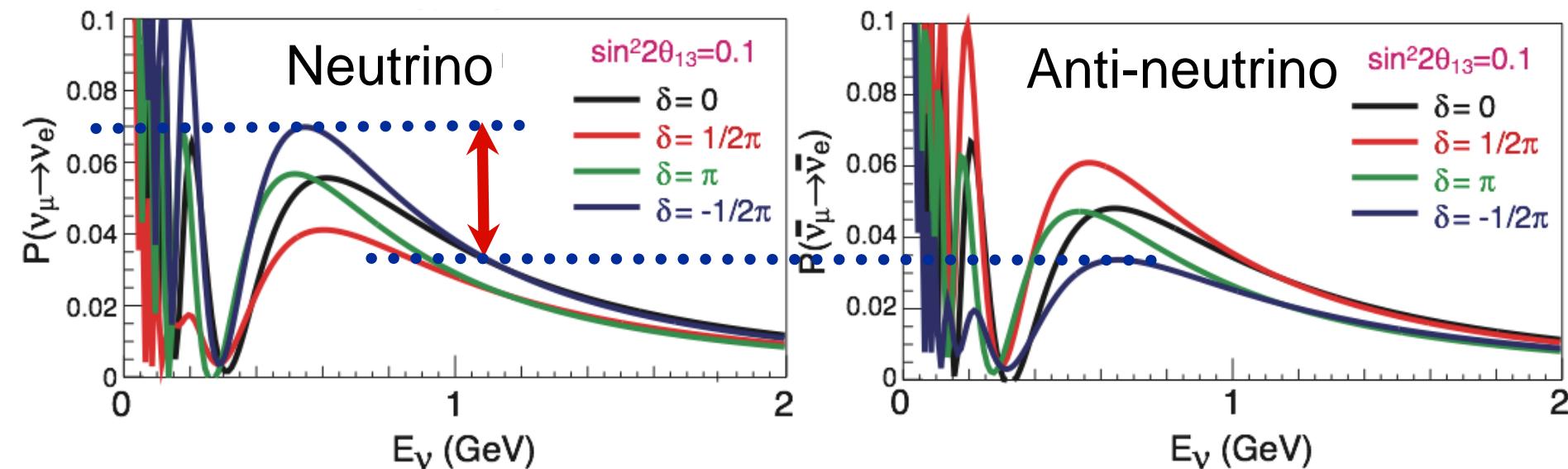




# CP measurements

$P(\nu_\mu \rightarrow \nu_e)$ :  $\nu_e$  appearance probability

for 295km baseline,  
normal hierarchy



- Comparison of  $P(\nu_\mu \rightarrow \nu_e)$  and  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$   
Max.  $\sim \pm 25\%$  ( $L=295$  km) change from  $\delta=0$  case
- Measure 1<sup>st</sup> and 2<sup>nd</sup> oscillation maxima in  $P(\nu_\mu \rightarrow \nu_e)$
- Comparison of accelerator  $P(\nu_\mu \rightarrow \nu_e)$  and reactor  $P(\bar{\nu}_e \rightarrow \bar{\nu}_e)$

Matter effect → fake CP violation, BUT sensitive instrument to determine mass hierarchy



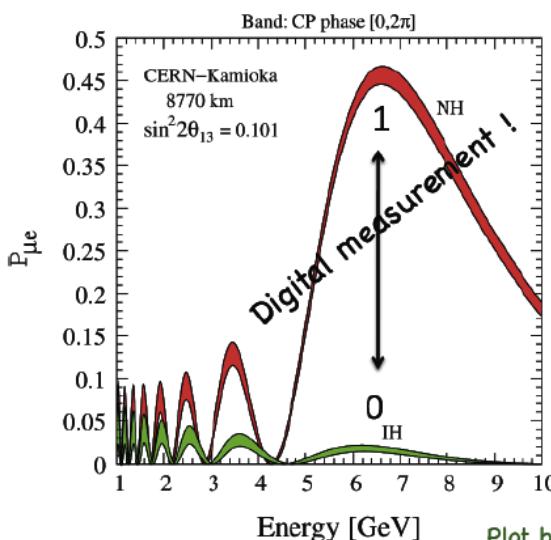
# Measurement of MH

Matter effect: appearance – accelerator experiments NOvA, LBNE, LBNO  
disappearance – PINGU, ORCA, INO, HyperK

Interference effect between  
solar and atmospheric oscillations: reactor experiments JUNO, RENO50

## accelerator experiments

MSW effect at ~6.5 GeV  
only neutrino beam

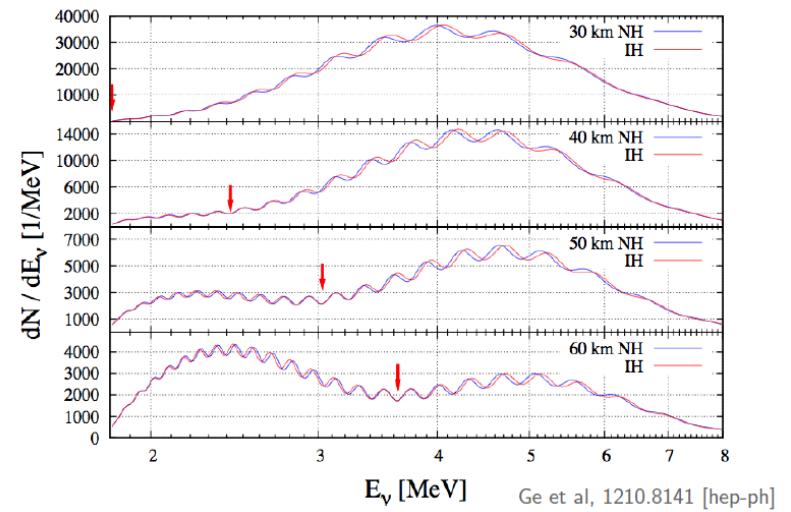


$$E_{\text{res}} \equiv \frac{\Delta m_{31}^2 \cos 2\theta_{13}}{2\sqrt{2}G_F n_e},$$

$$n_e(L)L|_{L_{\max}} = \frac{\pi}{\sqrt{2}G_F \tan 2\theta_{13}}$$

## reactor experiments

$$\begin{aligned} P_{ee} &= 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta_{21} \\ &\quad - \sin^2 2\theta_{13} \sin^2 \Delta_{31} \\ &\quad - \sin^2 \theta_{12} \sin^2 2\theta_{13} \sin^2 \Delta_{21} \cos 2\Delta_{31} \\ &\quad \pm \frac{1}{2} \sin^2 \theta_{12} \sin^2 2\theta_{13} \sin 2\Delta_{21} \sin 2\Delta_{31} \end{aligned}$$

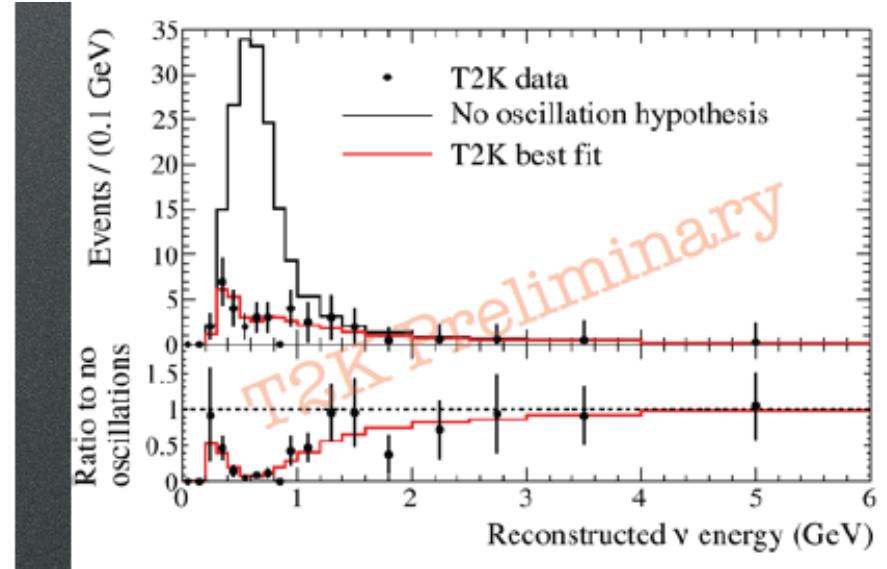
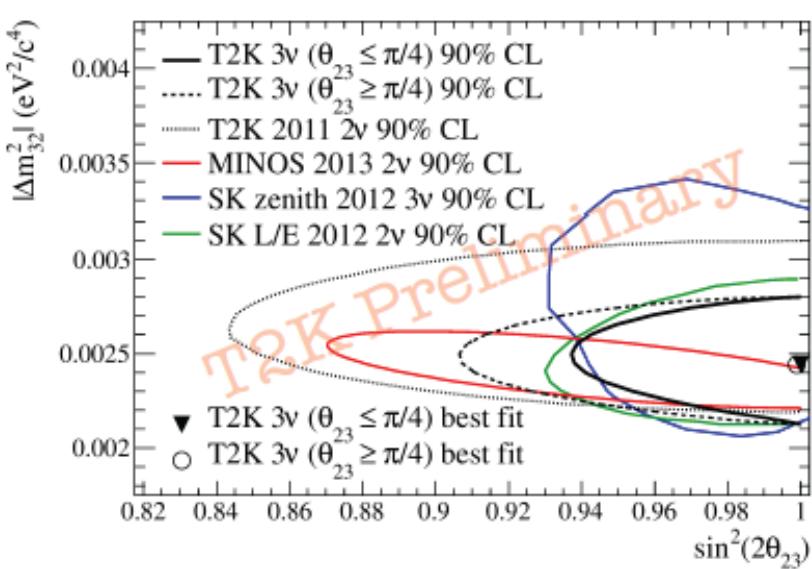




# $\theta_{23}$ measurement

$\nu_\mu$  disappearance

$$P(\nu_\mu \rightarrow \nu_\mu) \sim 1 - \left( \underbrace{\cos^4 \theta_{13} \cdot \sin^2 2\theta_{23}}_{\text{Leading}} + \underbrace{\sin^2 2\theta_{13} \cdot \sin^2 \theta_{23}}_{\text{Next-to-leading}} \right) \cdot \sin^2 \frac{\Delta m_{31}^2 \cdot L}{4E}$$

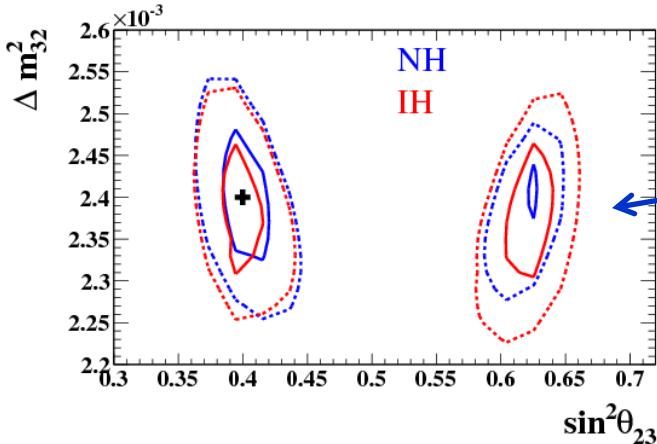




# Perspectives: $\theta_{23}$

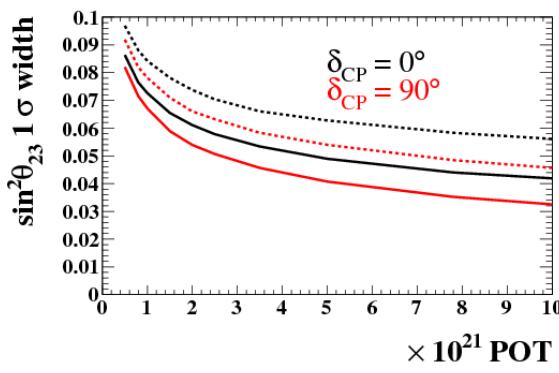
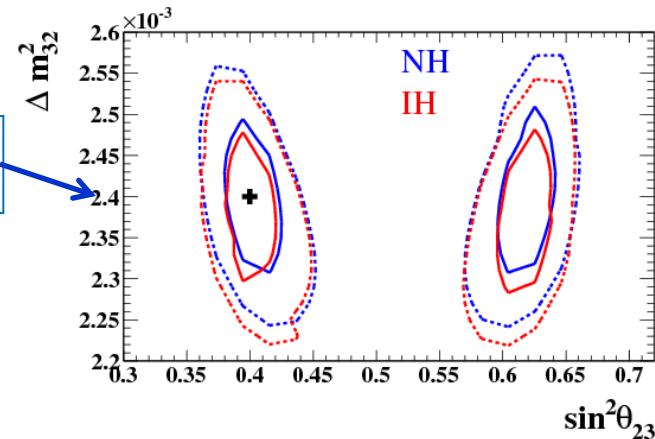
$7.8 \times 10^{21}$  pot  
 100%  $\nu$  running 90% CL  
 $\sin^2 2\theta_{13} = 0.1 +$  reactor data  
 solid: stat only; dashed: stat + current sys

T2K



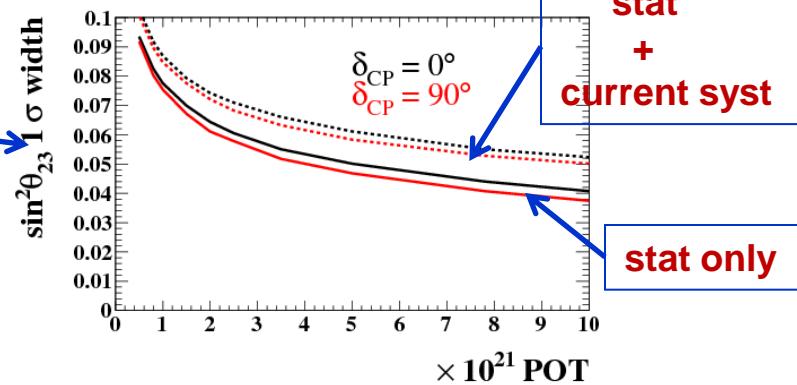
Octant

$7.8 \times 10^{21}$  pot  
 50%  $\nu$  - 50% anti- $\nu$  running 90% CL  
 $\sin^2 2\theta_{13} = 0.1 +$  reactor data  
 solid: stat only; dashed: stat + current sys



$\theta_{23}$

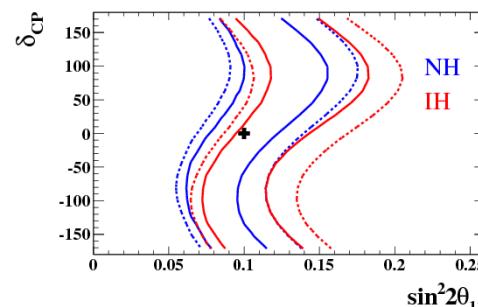
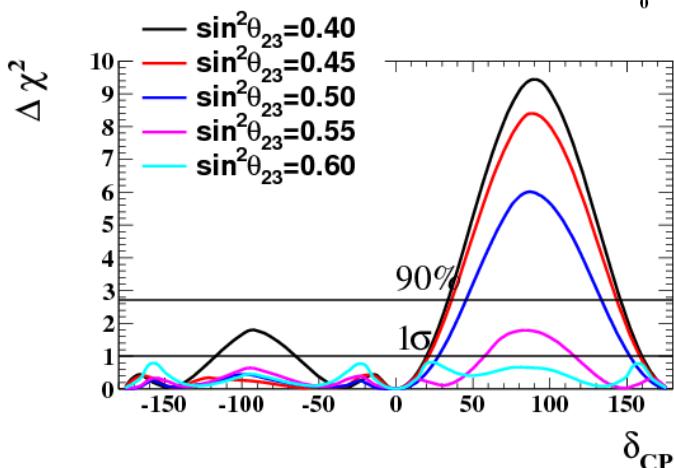
estimation  
 $\sigma(\theta_{23}) \sim 2$  deg





# T2K: sensitivity to CP

**IH**  $7.8 \times 10^{21}$  pot  
 100%  $\nu$  running  
 $\sin^2 2\theta_{13} = 0.1$   
 + reactor data



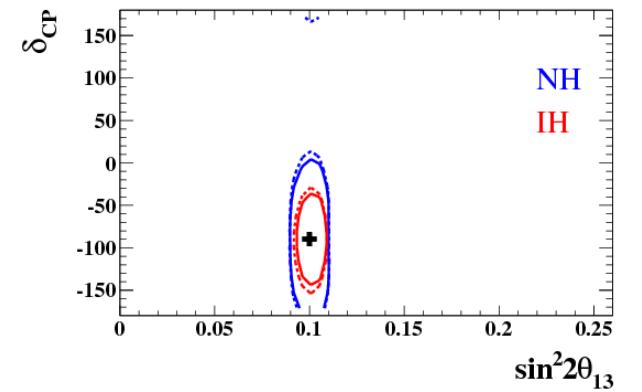
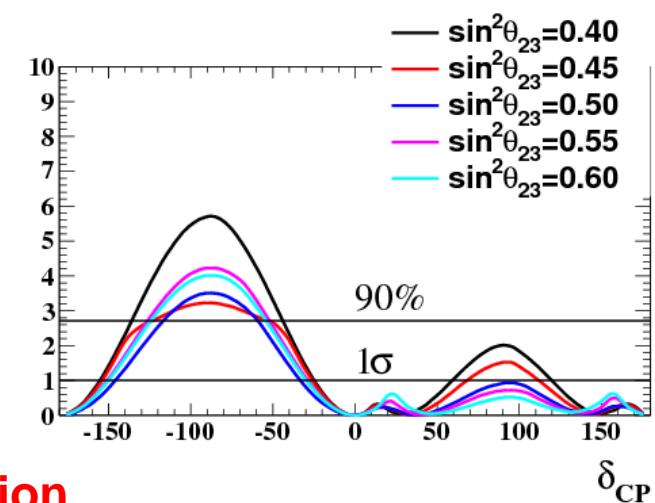
**T2K**

Chance to find an indication for CP violation  
 if  $\delta$  is about  $\pi/2$  or  $-\pi/2$

$7.8 \times 10^{21}$  pot  
 50%  $\nu$  - 50% anti-nu running  
 $\sin^2 2\theta_{13} = 0.1 +$  reactor data  
 $\sin^2 2\theta_{23} = 0.5$   
 $\delta = -90$  deg



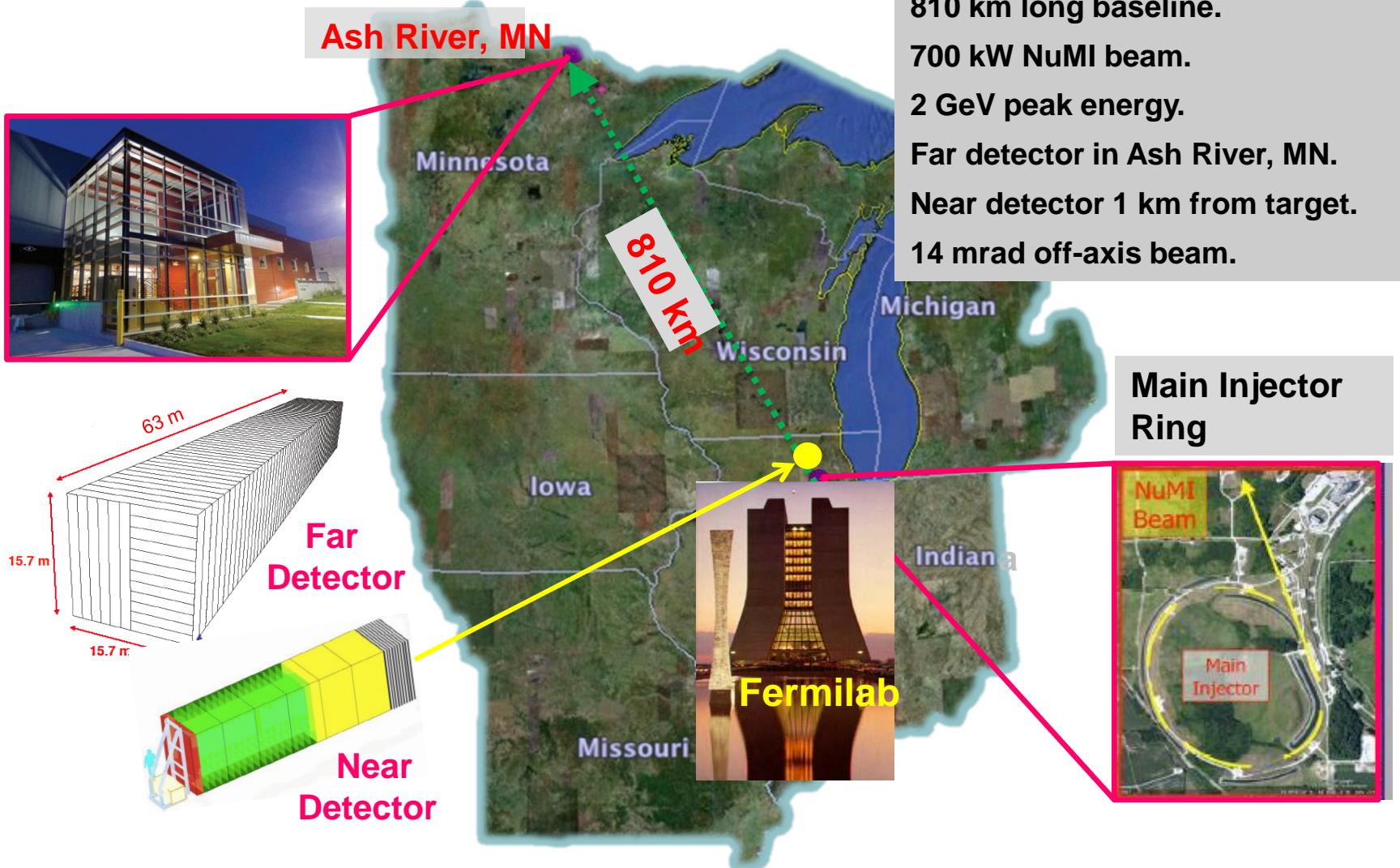
**NH**  $7.8 \times 10^{21}$  pot  
 50%  $\nu$  - 50% anti- $\nu$  running  
 $\sin^2 2\theta_{13} = 0.1 +$  reactor data





# Nova

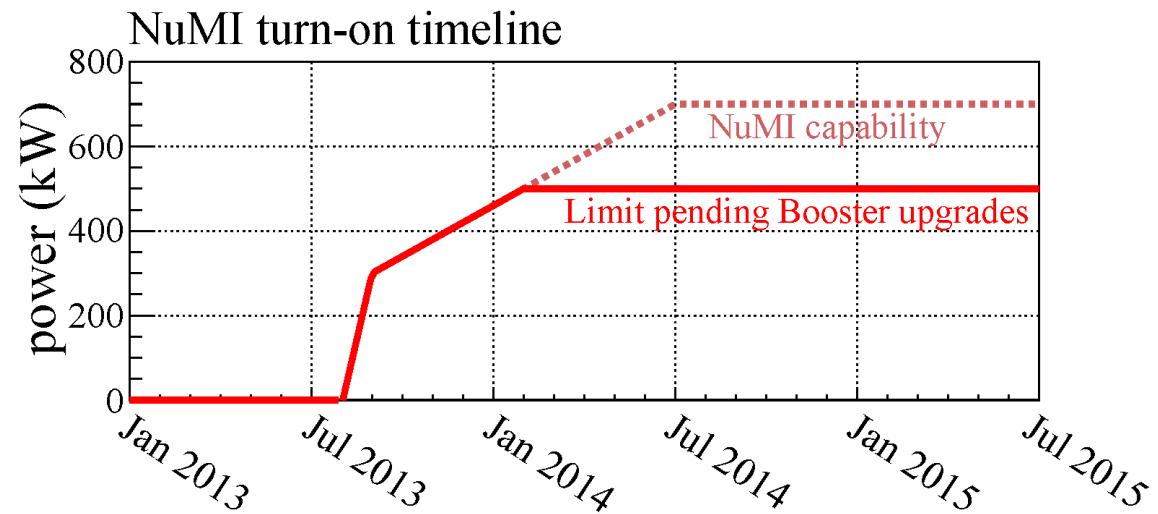
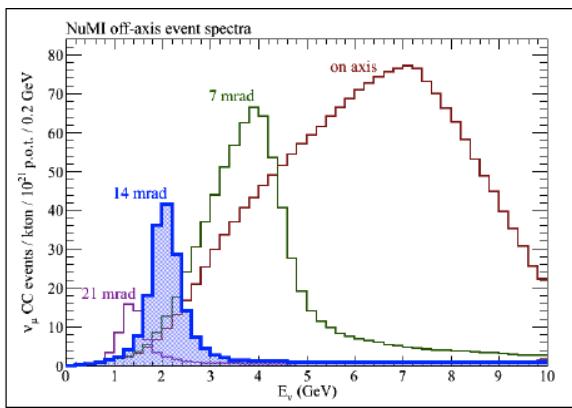
LBL experiment using neutrino off-axis  
narrow-band beam from FNAL





# Beam schedule

NOvA: first run in September 2013 with beam operating at  $\sim 300$  kW



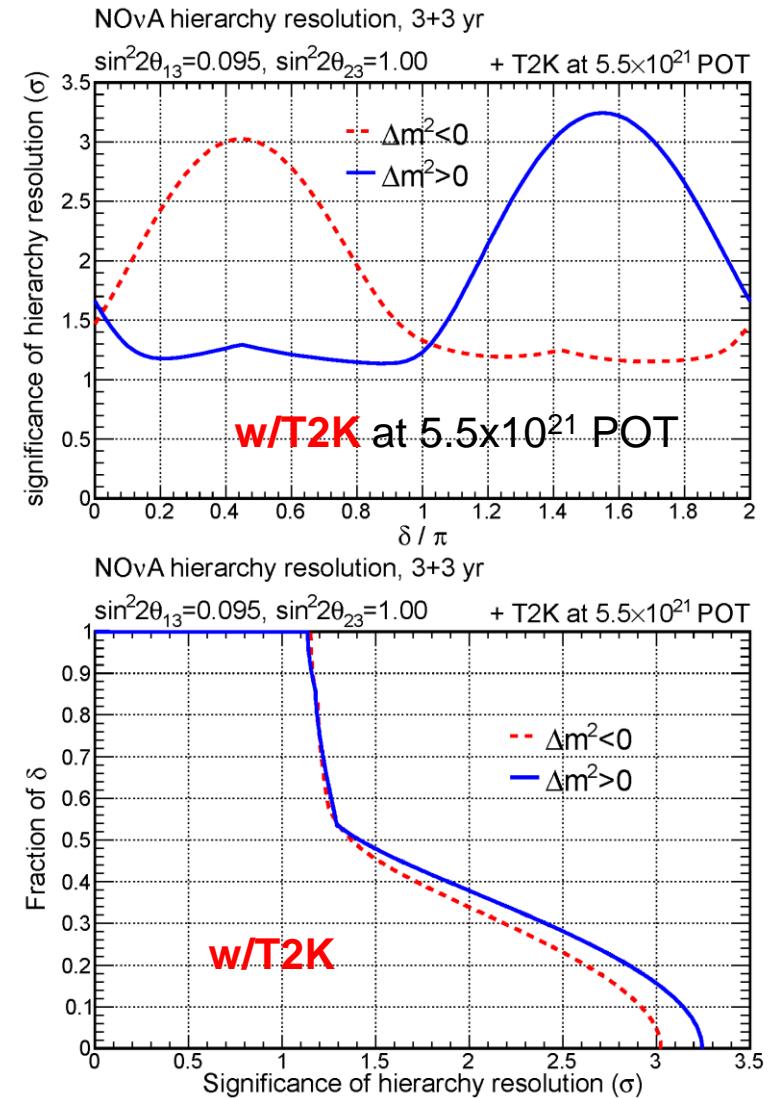
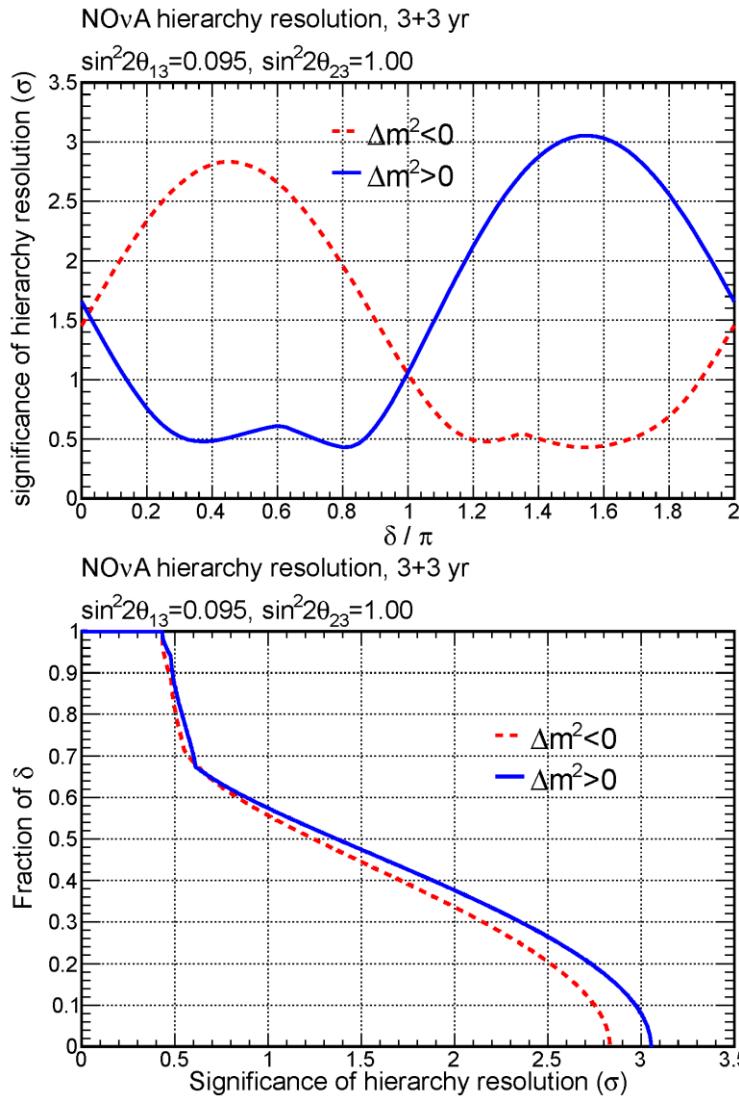
Beam intensity will be increased up to 500 kW next year and to 700 kW in 2 years.

Far Detector mass will be added at a rate of about 1 kton/3 weeks.  
Full installation of NOvA detectors will be completed in 2014.



# Mass Hierarchy

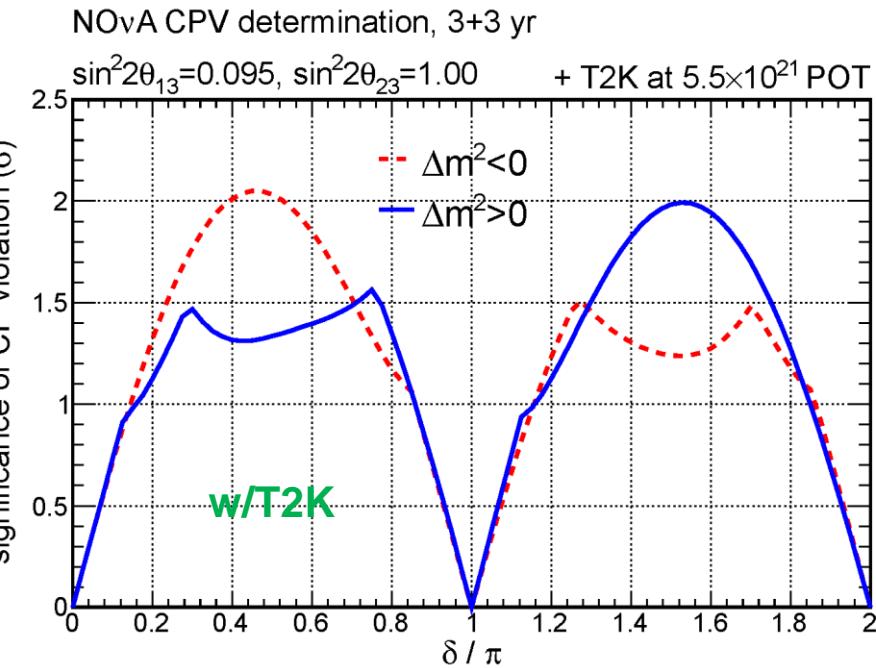
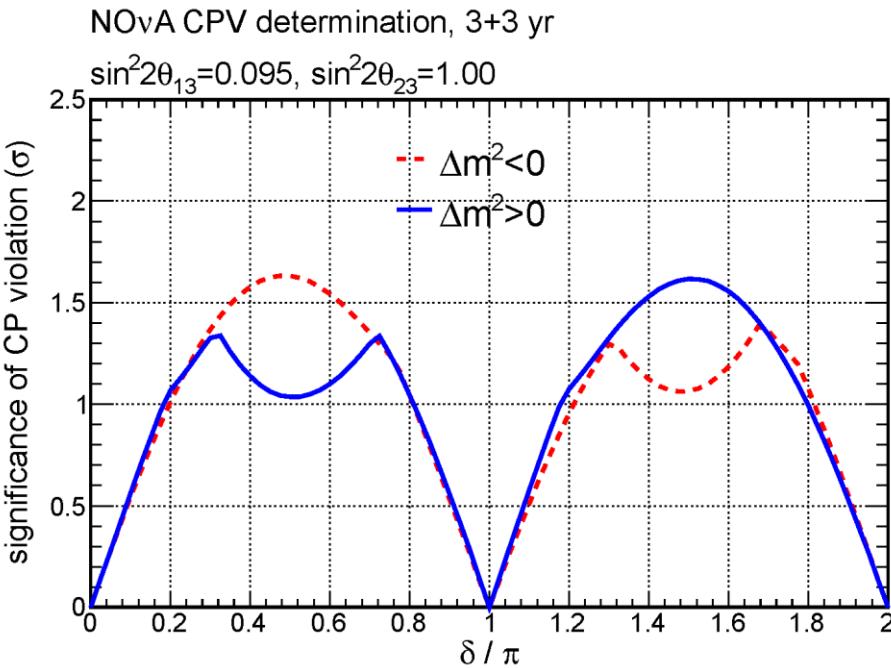
## Significance of the MH determination





# CP violation

Significance with which NOvA (+T2K) can establish CP violation.



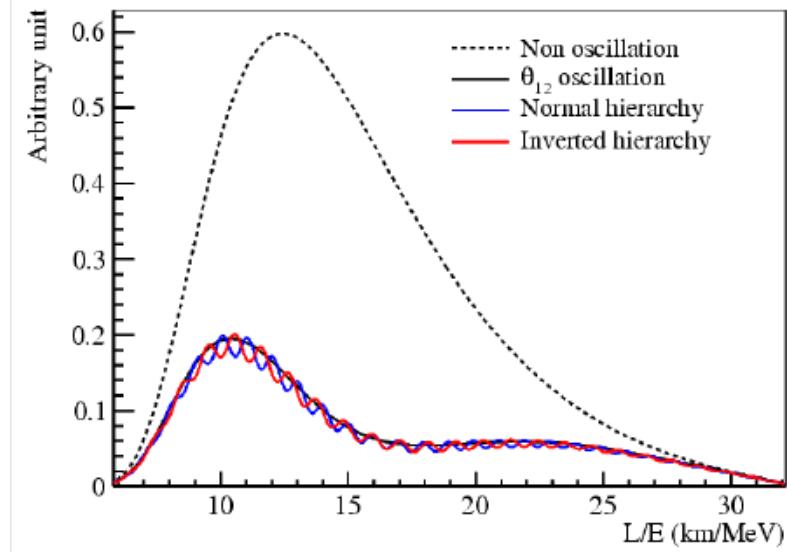
The significance goes to zero at  $\delta = 0$  and  $\delta = \pi$  since there is no CP violation at those points. The dips in the peaks occur because the mass ordering has not been resolved.

**Best case: CP violation at  $1.6\sigma$  (Nova only) and  $2.0\sigma$  (Nova + T2K)**

# **Far Future Prospects**



# JUNO and RENO50



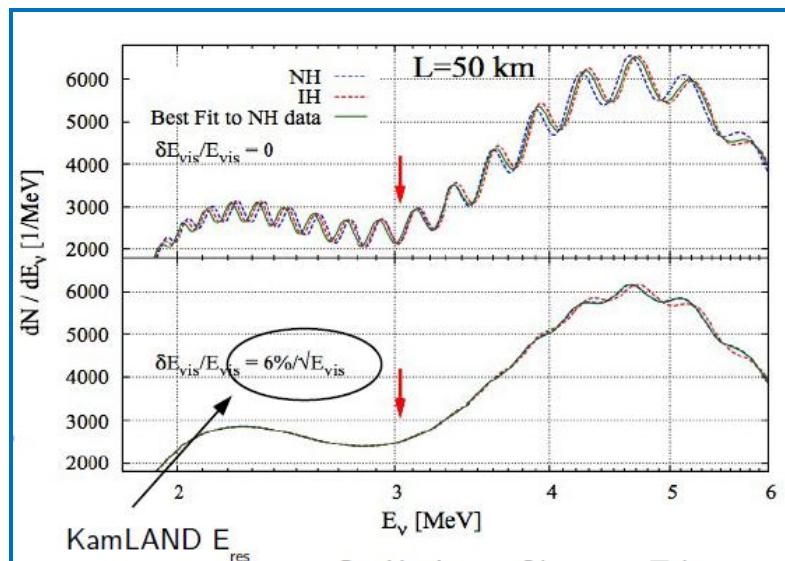
Two reactor experiments  
with  $L \sim 50$  km proposed  
in China and Korea

## RENO50

10 Kton liquid scintillator  
 $L \sim 47$  km (very close to optimal)  
RENO used as near detector  
Data taking expected to start in  
2019

## JUNO

20 Kton liquid scintillator  
 $L = 53$  km (very close to optimal)  
2 reactors, each of them with  
 $P \sim 18$  GW  
Data taking expected to start in  
2020



JUNO can determine the mass hierarchy  
with a confidence level of  $\Delta\chi^2 \sim 19$  ( $4.4\sigma$ )  
after 6 years of data taking



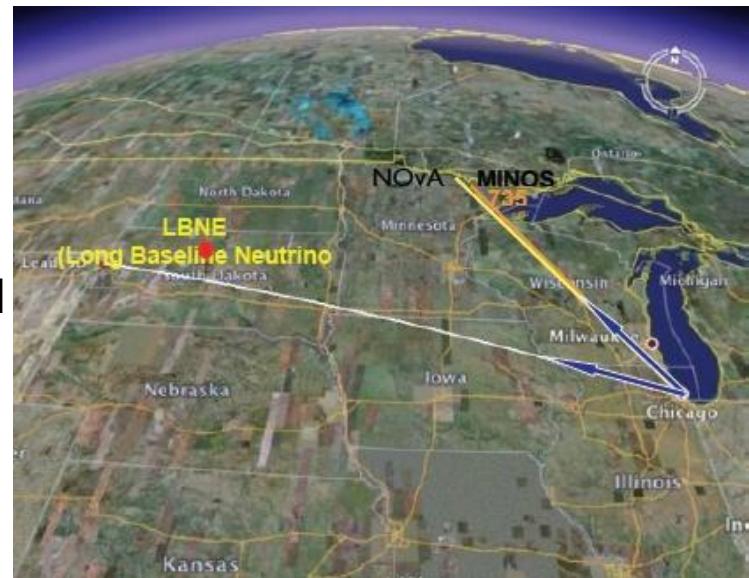
# LBNE

 $\nu_\mu \rightarrow \nu_e$ 

The US based LBL project

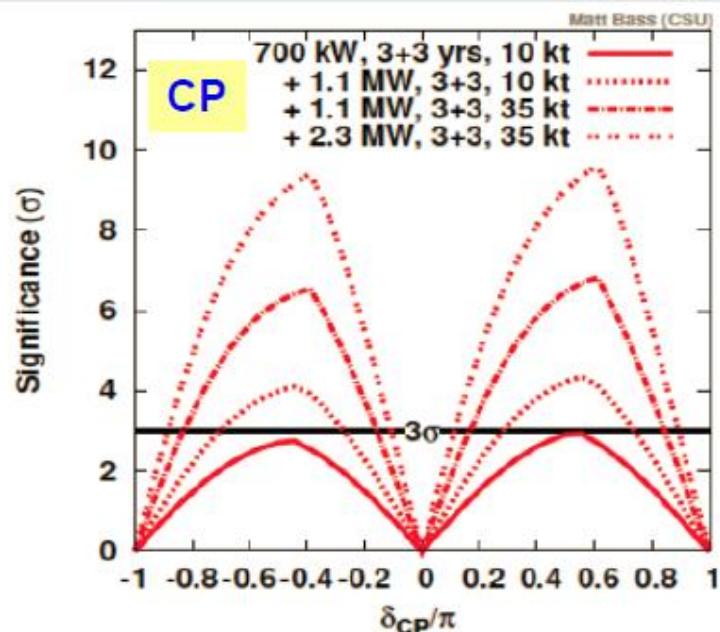
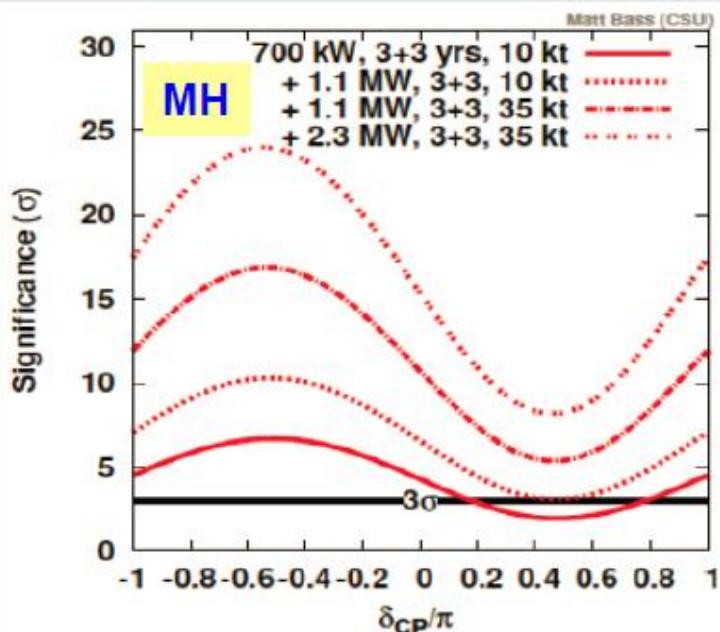
Neutrino beam from FNAL to Homestake  
L = 1300 km, Ep= 80-120 GeV, 0.7-2.3 kW NuMI  
 $E_\nu$  = 0.5 – 5 GeV

1<sup>st</sup> stage: far detector 10 kt LAr TPC, on surface  
2<sup>nd</sup> stage: far detector 34 kt Lar TPC, underground



Sensitivity to MH and CP phase

M.Diwan, NeuTel 2013





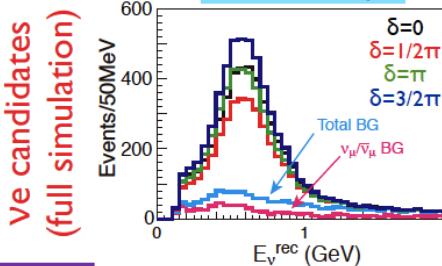
## The LBL project in Japan

- JPARC 1.66MW
- off-axis  $\nu$  beam
- $L = 296$  km
- far detector HyperKamiokande total mass 1 Mt

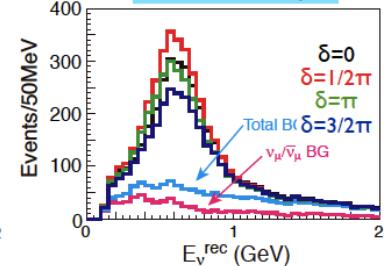
### Example: case of Hyper-K

$$\sin^2 2\theta_{13} = 0.1$$

$\nu$  0.75MW × 3yrs



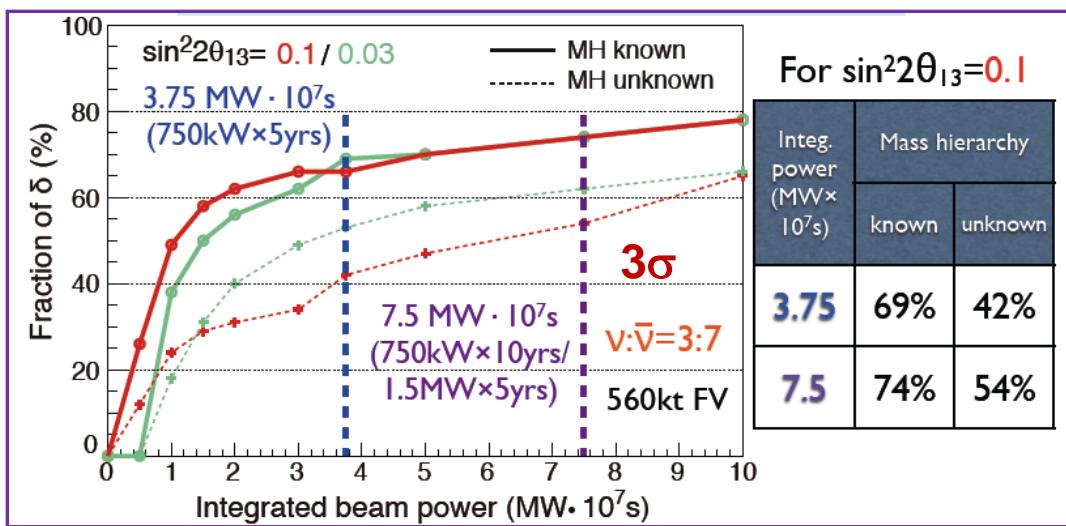
$\nu$  0.75MW × 7yrs



Expected number of  $\nu_e$  candidate events for  $\delta=0$

$\nu_\mu \rightarrow \nu_e$ CC	3606
$\nu_\mu + \bar{\nu}_\mu$ CC	35
$\nu_e + \bar{\nu}_e$ CC	880
NC	649

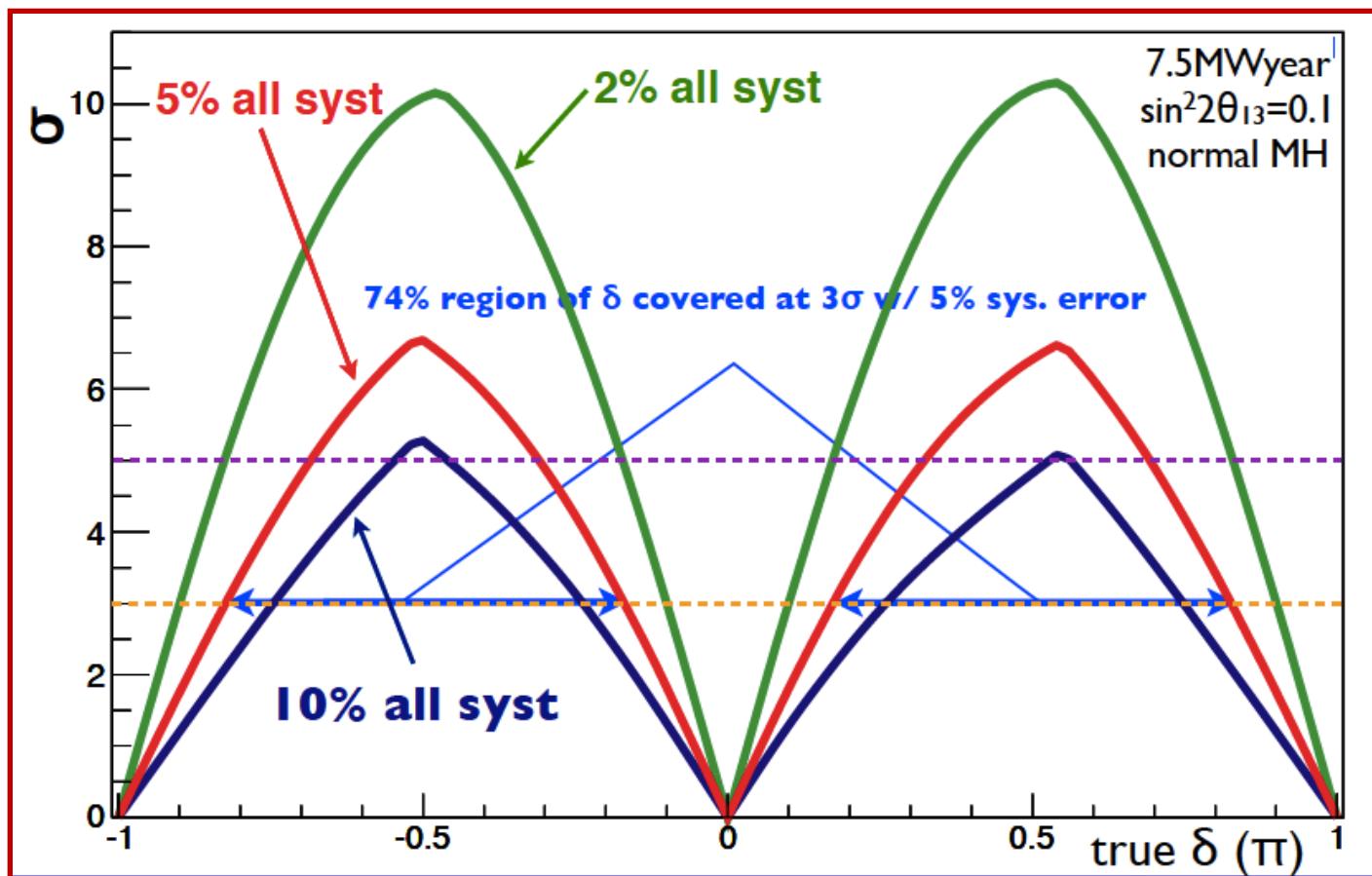
$\nu_\mu \rightarrow \nu_e$ CC	2339
$\nu_\mu + \bar{\nu}_\mu$ CC	23
$\nu_e + \bar{\nu}_e$ CC	878
NC	678





# T2HK: CPV discovery potential

MH is known !

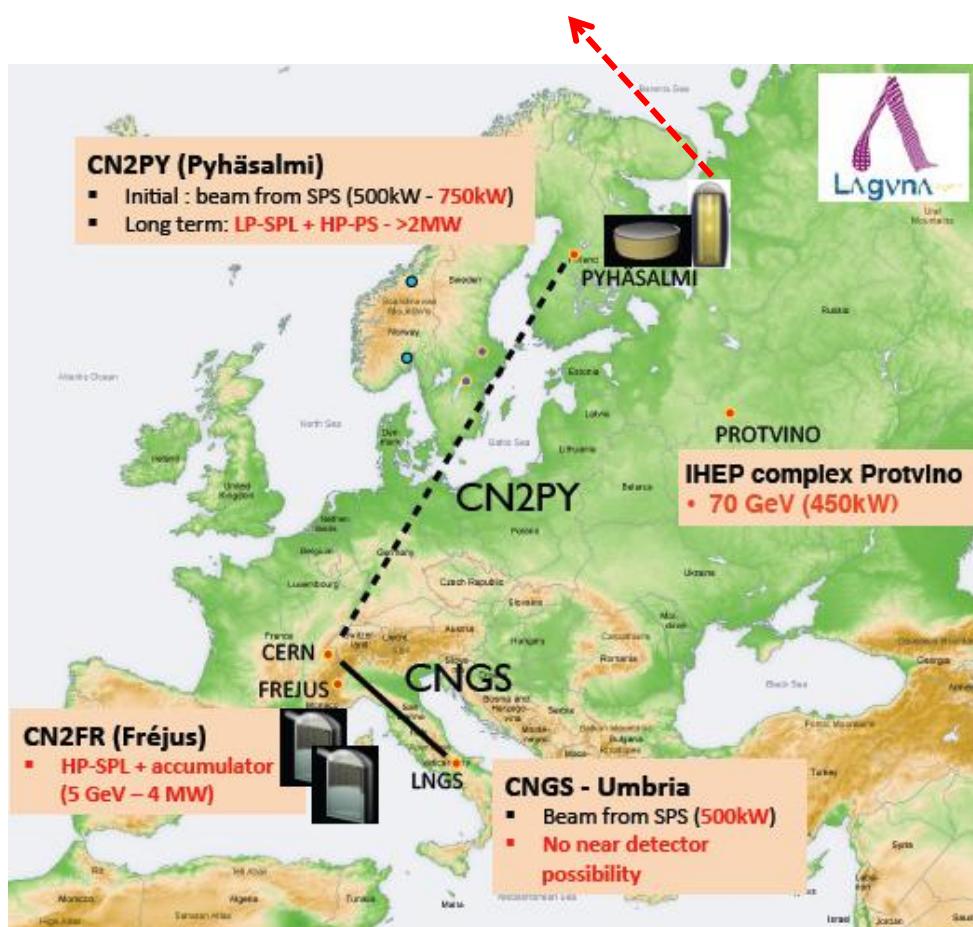


High sensitivity to CP phase for systematics < 5%

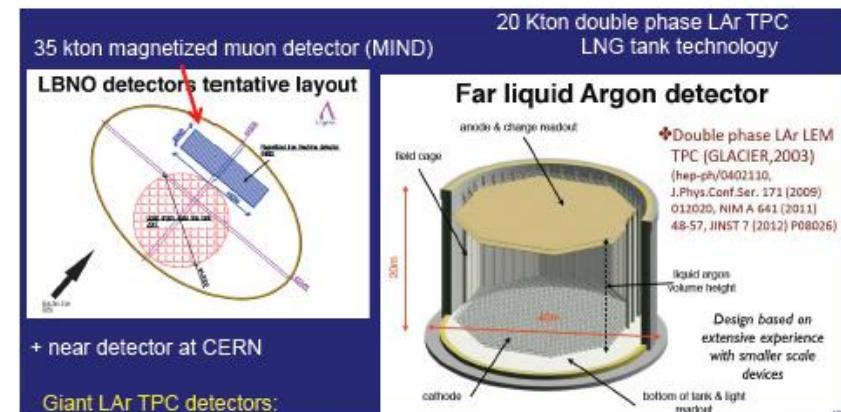


# LAGUNA/LBNO

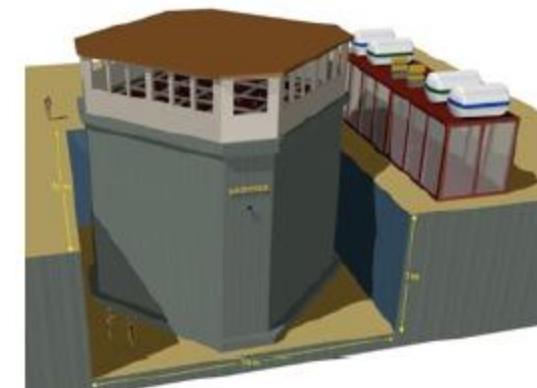
European LBL project  
Wide band neutrino beam  
from CERN to Pyhasalmi (Finland)  
 $L = 2300$  km



## Far detector

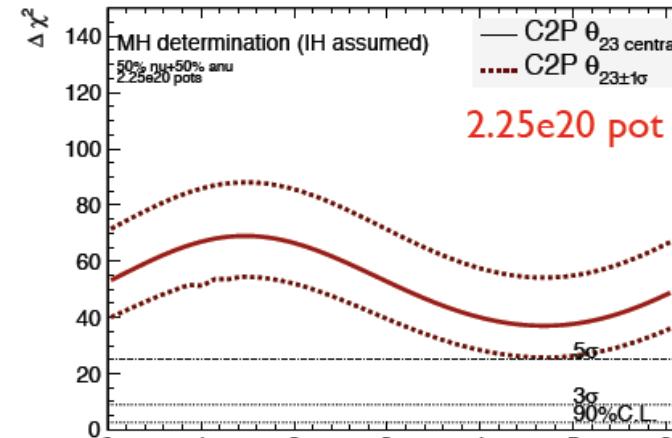
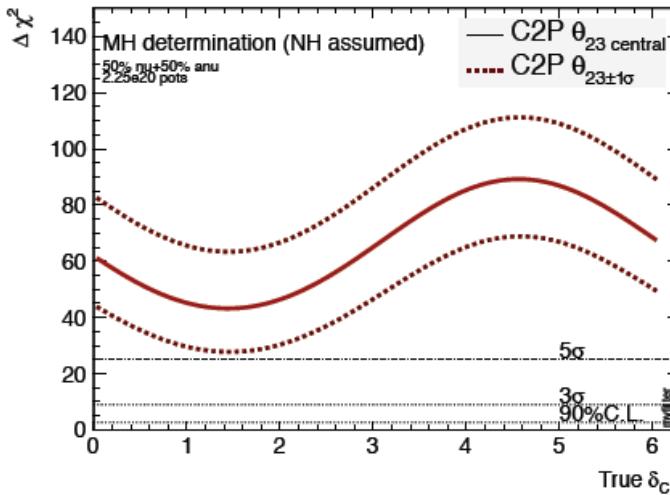


## R&D: LAr demonstrator at CERN



# LAGUNA/LBNO sensitivity

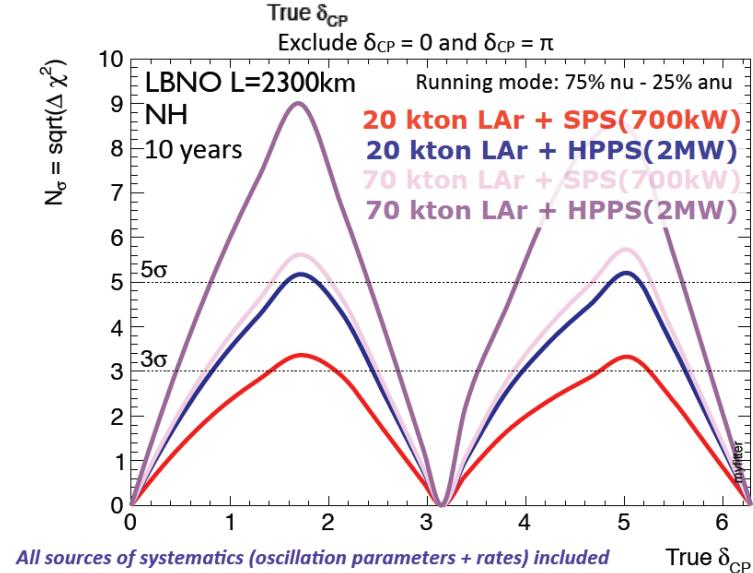
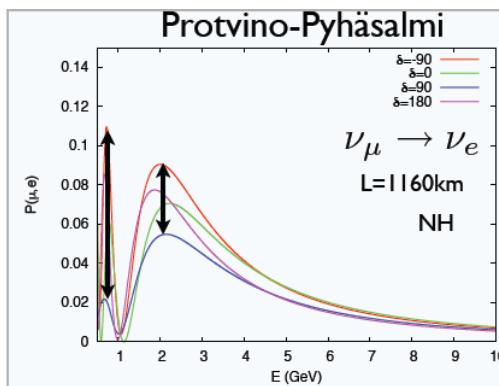
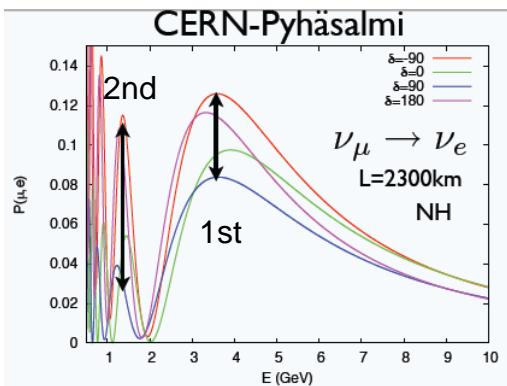
More than  $5\sigma$  determination of MH for all  $\delta$  values



A.Rubbia,  
talk at Lomonosov  
Conference 2013

Sensitivity to  $\delta$

Measurement in 1<sup>st</sup> and 2<sup>nd</sup> oscillation maxima





# Conclusion

- 2011-2013

- $\theta_{13}$  is measured and large

*Open very exiting perspectives in neutrino oscillations*

- Observation of  $\nu_\mu \rightarrow \nu_e$  appearance at  $7.5\sigma$  significance

*A new type of transformation among neutrinos has firmly established*

- Near future:

- precision measurements of neutrino mixing parameters
  - an initial search for CP violation in lepton sector

- Far future:

- measurement of CP violation in lepton sector
  - determination of neutrino mass hierarchy