

# Generalized metrics in the presence of a scalar field

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A. Azizallahi, B. M, A. Hajibarat, H. Anjomshoa, arXiv: 2307.09328

B. M, P.K Kangazi, F. Sadeghi, arXiv: 2307.13588

# Introduction

## Action:

$$S = \int d^4x (R - \partial_\mu \varphi(r) \partial^\mu \varphi(r)).$$

## Equations of motion:

$$G_{\mu\nu} = \partial_\mu \varphi(r) \partial_\nu \varphi(r) - \frac{1}{2} g_{\mu\nu} \partial_\beta \varphi(r) \partial^\beta \varphi(r),$$

$$R_{\mu\nu} = \partial_\mu \varphi(r) \partial_\nu \varphi(r),$$

$$\nabla_\mu \nabla^\mu \varphi(r) = 0.$$

# Introduction

**FJNW metric:**

$$ds^2 = -f^\gamma dt^2 + f^{-\gamma} dr^2 + f^{1-\gamma} r^2 (d\theta^2 + \sin^2 \theta d\phi^2),$$

**Metric function:**

$$f(r) = 1 - \frac{2m}{r}.$$

This metric is not the most general solution for the action.

The metric has singularities at the following points:

$$r_{singularity} = \begin{cases} 0 \\ 2m \end{cases}$$

# Generalised three parameter metrics

**The generalised three parameter metrics:**

$$ds^2 = -f^\gamma dt^2 + f^\mu k^\nu \left( \frac{dr^2}{f} + r^2 d\theta^2 \right) + f^\beta r^2 d\phi^2,$$

**Metrics functions:**

$$f(r) = 1 - \frac{2m}{r},$$

$$k(r, \theta) = 1 - \frac{2m}{r} + \frac{m^2}{r^2} N^2(\theta).$$

## Generalised three parameter metrics

**From**  $R_{\theta\theta} = 0$ , **to the first order of m**  $\Rightarrow \beta = 1 - \gamma.$

**From**  $R_{r\theta} = 0$ , **to the first order of m**  $\Rightarrow \mu + \nu = \beta.$

$$\mu + \nu = 1 - \gamma.$$

**From**  $R_{r\theta} = 0$ , **to the second order of m**  $\Rightarrow \mathcal{N}(\theta) = \sin \theta.$

$$R_{rr} = \partial_r \varphi(r) \partial_r \varphi(r), \quad \Rightarrow \quad \varphi(r) = \sqrt{\frac{1 - \gamma^2 - \nu}{2}} \ln \left(1 - \frac{2m}{r}\right).$$

# Generalised three parameter metrics

**Metrics:**

$$ds^2 = -f^\gamma dt^2 + f^\mu k^\nu \left( \frac{dr^2}{f} + r^2 d\theta^2 \right) + f^{1-\gamma} r^2 \sin^2 \theta d\phi^2,$$

$$\mu + \nu = 1 - \gamma,$$

**Metrics functions:**

$$f(r) = 1 - \frac{2m}{r}, \quad k(r, \theta) = 1 - \frac{2m}{r} + \frac{m^2 \sin^2 \theta}{r^2}.$$

**parameters limits:**

$$\nu = 0, \quad \mu = 1 - \gamma \quad \Rightarrow \quad \text{FJNW metric}$$

$$\nu = 1 - \gamma^2, \quad \mu = \gamma^2 - \gamma, \quad \varphi(r) = 0 \quad \Rightarrow \quad \gamma\text{-metric}$$

# Generalised three parameter metrics

**The four-dimensional  $\gamma$ -metric in isotropic form:**

$$ds^2 = -f^\gamma dt^2 + \left(1 + \frac{m}{2\rho}\right)^4 [f^\mu k^\nu (d\rho^2 + \rho^2 d\theta^2) + \rho^2 f^\beta \sin^2 \theta d\phi^2],$$

$$r = \rho \left(1 + \frac{m}{2\rho}\right)^2,$$

**Metrics parameters:**

$$\mu + \nu = 1 - \gamma,$$

**Metrics functions:**

$$f = \left(\frac{m - 2\rho}{m + 2\rho}\right)^2, \quad k = \frac{m^4 + 16\rho^4 - 8m^2\rho^2 \cos(2\theta)}{(m + 2\rho)^4}.$$

# Generalised three parameter metrics

Ricci scalar:

$$R = \frac{2(\gamma^2 + \nu - 1)m^2}{r^{2+\gamma-\nu}(r - 2m)^{2-\gamma-\nu}(r^2 - 2mr + m^2 - m^2 \cos^2 \theta)^\nu}$$

Singularities points:

$$r_{singularity} = \begin{cases} 0 & \text{for } \nu - \gamma < 2 \\ 2m & \text{for } \nu + \gamma < 2 \\ m(1 \pm \cos \theta) & \text{for } \nu > 0 \end{cases}$$

# Generalised three parameter metrics

## Five dimensional metrics:

$$ds^2 = -f^\gamma dt^2 + f^\mu k^\nu \left( \frac{dr^2}{f} + r^2 d\theta^2 \right) + r^2 f^\beta (\sin^2 \theta d\phi^2 + \cos^2 \theta d\psi^2),$$
$$\mu + \nu = \beta, \quad \beta = \frac{1 - \gamma}{2}.$$

Higher dimensions?

## Metrics functions:

$$f(r) = 1 - \frac{2m}{r^2}, \quad k(r, \theta) = 1 - \frac{2m}{r^2} + \frac{m^2 \sin^2(2\theta)}{r^4},$$

$$R_{rr} = \partial_r \varphi(r) \partial_r \varphi(r) \quad \Rightarrow \quad \varphi(r) = \sqrt{\frac{\frac{3}{4}(1 - \gamma^2) - \nu}{2}} \ln\left(1 - \frac{2m}{r^2}\right).$$

# Generalised three parameter metrics

## Metrics limits:

five-dimensional JNW metric:

$$\nu = 0, \quad \mu = \frac{1 - \gamma}{2}$$

five-dimensional  $\gamma$ -metric:

$$\nu = \frac{3}{4} (1 - \gamma^2), \quad \mu = \frac{1}{4} (3\gamma^2 - 2\gamma - 1), \quad \varphi(r) = 0$$



# Generalised three parameter metrics

Ricci scalar in five dimensions:

$$R = \frac{2 m^2 (2 \gamma + 4 \mu + 1 - 3 \gamma^2)}{r^{2+\gamma-2\nu+3} (r^2 - 2 m)^{\frac{3-\gamma-2\nu}{2}} (r^4 - 2 m r^2 + m^2 \sin^2(2\theta))^\nu}$$

Singularities points:

$$r_{singularity} = \begin{cases} 0 & \text{for } 2\nu - \gamma < 3 \\ \sqrt{2m} & \text{for } \gamma + 2\nu < 3 \\ \sqrt{m(1 \pm \cos(2\theta))} & \text{for } \nu > 0 \end{cases}$$



# Generalised three parameter metrics

The five-dimensional  $\gamma$ -metric in isotropic form:

$$ds^2 = -f^\gamma dt^2 + \left(1 + \frac{m}{2\rho}\right)^2 \left[ f^\mu k^\nu (d\rho^2 + \rho^2 d\theta^2) + \rho^2 f^\beta (\sin^2 \theta d\phi^2 + \cos^2 \theta d\psi^2) \right],$$
$$r = \rho \left(1 + \frac{m}{2\rho^2}\right),$$

Metrics parameters:

$$\mu + \nu = \frac{1 - \gamma}{2}$$

Metrics functions:

$$f = \left(\frac{m - 2\rho^2}{m + 2\rho^2}\right)^2, \quad k = \frac{m^4 + 16\rho^8 - 8m^2\rho^4 \cos(4\theta)}{(m + 2\rho^2)^4}.$$

# Generalised three parameter metrics

**Three parameter metrics in toroidal coordinates:**

$$ds^2 = -f^\gamma dt^2 + \frac{f^\mu}{f} dr^2 + r^2 f^\mu d\theta^2 + r^2 f^{\frac{1-\gamma}{d-3}} \sum_{i=1}^{d-3} d\phi_i^2,$$

**Metrics function:**

$$f(r) = \frac{2m}{r^{d-3}}.$$

**Scalar field:**

$$\varphi(r) = \sqrt{\frac{d-2\gamma-(d-2)\gamma^2}{4(d-3)}} - \frac{\mu}{2} \ln\left(\frac{2m}{r^{d-3}}\right).$$

# Generalised three parameter metrics

**Three parameter metrics in flat coordinates:**

$$ds^2 = -f^\gamma dt^2 + \frac{f^\mu k^\nu}{f} dr^2 + r^2 f^\mu k^\nu d\theta^2 + \theta^2 r^2 f^{\frac{1-\gamma}{d-3}} d\phi^2$$
$$+ r^2 f^{\frac{1-\gamma}{d-3}} \sum_{i=1}^{d-4} d\psi_i^2, \quad \mu + \nu = \frac{1-\gamma}{d-3},$$

**Metrics functions:**

$$f(r) = \frac{2m}{r^{d-3}}, \quad k(r, \theta) = \frac{2m}{r^{d-3}} + \frac{(d-3)^2 m^2 \theta^2}{r^{2(d-3)}}.$$

**Scalar field:**

$$\varphi(r) = \sqrt{\frac{(d-2)(1-\gamma^2)}{4(d-3)} - \frac{\nu}{2}} \ln\left(\frac{2m}{r^{d-3}}\right).$$

# Rotating three parameter metrics

**Lewis-Weyl-Papapetrou (LWP) metric:**

$$ds^2 = -f(dt - \omega d\phi)^2 + f^{-1} [e^{2\eta} (d\rho^2 + dz^2) + \rho^2 d\phi^2].$$

**Coordinate transformations:**

$$\rho = (x^2 - 1)^{\frac{1}{2}} (1 - y^2)^{\frac{1}{2}},$$

$$z = xy.$$

**LWP metric in  $x$  and  $y$  coordinates:**

$$ds^2 = -f(dt - \omega d\phi)^2 + \frac{\sigma^2}{f} [e^{2\eta} (x^2 - y^2) \left( \frac{dx^2}{x^2 - 1} + \frac{dy^2}{1 - y^2} \right) + (x^2 - 1)(1 - y^2) d\phi^2].$$

## Rotating three parameter metrics

**From  $R_{tt}$  and  $R_{t\phi}$ :**

$$\partial_x \omega = \sigma(1 - y^2) f^{-2} \partial_y u, \quad \partial_y \omega = \sigma(1 - x^2) f^{-2} \partial_x u.$$

**From  $R_{xx}$ ,  $R_{yy}$  and  $R_{xy}$ :**

$$\partial_x \eta = \dots + 4\pi \frac{(1 - y^2)}{(x^2 - y^2)} x(x^2 - 1)(\partial_x \varphi)^2,$$

$$\partial_y \eta = \dots + 4\pi \frac{(x^2 - 1)^2 y}{(x^2 - y^2)} (\partial_x \varphi)^2.$$

**From  $\partial_y \partial_x \eta = \partial_x \partial_y \eta$ :**

$$\varphi'' + \frac{2x}{x^2 - 1} \varphi' = 0, \quad \Rightarrow \quad \varphi = c_1 \ln\left(\frac{x-1}{x+1}\right) + c_2.$$

# Rotating three parameter metrics

**Ernst equations:**

$$(\varepsilon + \varepsilon^*) \Delta \varepsilon = 2 (\nabla \varepsilon)^2,$$

$$\Delta \equiv \sigma^{-2} (x^2 - y^2)^{-1} \partial_x [(x^2 - 1) \partial_x] + \partial_x [(x^2 - 1) \partial_x],$$

$$\nabla \equiv \sigma^{-1} (x^2 - y^2)^{-\frac{1}{2}} \hat{i} [(x^2 - 1)^{\frac{1}{2}} \partial_x] + \hat{j} [(1 - y^2)^{\frac{1}{2}} \partial_x].$$

**Ernst potential:**

$$\varepsilon = f + i u,$$

$$\sigma = \sqrt{m^2 - a^2},$$

$a$  = rotation parameter.

## Rotating three parameter metrics

**Ernst potential:**

$$\varepsilon = \left( \frac{x-1}{x+1} \right)^q \frac{x-1 + (x^2-1)^{-q} d_+}{x+1 + (x^2-1)^{-q} d_-},$$

$$d_{\pm} = -\alpha^2 (x \pm 1) h_+ h_- (x^2 - 1)^{-q}$$

$$+ i\alpha [y(h_+ + h_-) \pm (h_+ - h_-)],$$

$$h_{\pm} = (x \pm y)^{2q},$$

$$\alpha = -\frac{a}{m + \sigma},$$

$$q = 1 - \gamma.$$

# Rotating three parameter metrics

Metrics functions :

$$f = \frac{A}{B},$$

$$\omega = -2 \left( a + \sigma \left( \frac{C}{A} \right) \right),$$

$$e^{2\eta} = \frac{1}{4} \left( 1 + \frac{m}{\sigma} \right)^2 \frac{A}{(x^2 - 1)^{1+q}} \left( \frac{x^2 - 1}{x^2 - y^2} \right)^{1-\nu},$$

$$A = a_+ a_- + b_+ b_-, \quad B = a_+^2 + b_+^2,$$

$$C = (x+1)^q [x(1-y^2)(\lambda + \xi)a_+ + y(x^2-1)(1-\lambda\xi)b_+],$$

$$a_{\pm} = (x \pm 1)^q [x(1 - \lambda\xi) \pm (1 + \lambda\xi)],$$

$$b_{\pm} = (x \pm 1)^q [y(\lambda + \xi) \mp (\lambda - \xi)],$$

$$\lambda = \alpha (x^2 - 1)^{-q} (x + y)^{2q}, \quad \xi = \alpha (x^2 - 1)^{-q} (x - y)^{2q}.$$

# Rotating three parameter metrics

**Coordinate conversion:**

$$x = \frac{r - m}{\sqrt{m^2 - a^2}}, \quad y = \cos\theta.$$

**Scalar field:**

$$\square\varphi = 0, \quad \Rightarrow \quad \varphi(r) = \sqrt{\frac{1 - \gamma^2 - \nu}{2}} \ln \left( \frac{r - m - \sqrt{m^2 - a^2}}{r - m + \sqrt{m^2 - a^2}} \right).$$

**Ricci tensor:**

$$R_{\mu\nu} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & \partial_r \varphi \partial_r \varphi & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}.$$

# Rotating three parameter metrics

## Limits of metrics:

$$\nu = 1 - \gamma^2, \quad \mu = \gamma^2 - \gamma \quad \Rightarrow \quad \text{rotating } \gamma\text{-metric}$$

$$\nu = -\mu, \quad \gamma = 1 (q = 0) \quad \Rightarrow \quad \text{Bogush-Gal'tsov (BG) metric}$$

$$\nu = 1 - \gamma = -q, \quad \mu = 0 \quad \Rightarrow \quad \text{rotating metric at } \nu = -q$$

$$\nu = 0, \quad \mu = 1 - \gamma \quad \Rightarrow \quad \text{rotating FJNW metric}$$



I. Bogush, D. Gal'tsov, arXiv: 2001.02936

B. M, P.K Kangazian Kangazi, F. Sadeghi, arXiv: 2307.13588

## Rotating three parameter metrics

**Quasi Normal Modes (QNMs) for rotating JNW metric  
by using light-ring method**  $\nu = 0$ ,  $\mu = 1 - \gamma$ :

$$\begin{aligned} Q_{JNW} &= (\Omega + i\Gamma)_{JNW} \\ &= j \left[ \frac{1}{3\sqrt{3}m} \pm \frac{1}{3\sqrt{3}m} \left( \frac{11}{54} \left( \frac{a}{m} \right)^2 + \frac{2\sqrt{3}a}{9m} \right. \right. \\ &\quad \left. \left. - q \ln 3 \right) \right] + i \left( n + \frac{1}{2} \right) \left\{ \frac{1}{3\sqrt{3}m} \left[ 1 - \frac{2a^2}{27m^2} \right. \right. \\ &\quad \left. \left. - q \ln 3 \right] \right\}. \end{aligned}$$

## Rotating three parameter metrics

**QNMs for BG metric**  $\gamma = 0, \mu = -\nu$ :

$$\begin{aligned} Q_{BG} &= (\Omega + i\Gamma)_{BG} \\ &= j \left[ \frac{1}{3\sqrt{3}m} \pm \frac{1}{3\sqrt{3}m} \left( \frac{11}{54} \left( \frac{a}{m} \right)^2 + \frac{2\sqrt{3}a}{9m} \right) \right] \\ &\quad + i \left( n + \frac{1}{2} \right) \left\{ \frac{1}{3\sqrt{3}m} \left[ 1 - \frac{2a^2}{27m^2} - \nu(\ln 2 \right. \right. \\ &\quad \left. \left. - \frac{\ln 3}{2} \right) \right] \right\}. \end{aligned}$$

## Rotating three parameter metrics

**Comparison between stability of rotating JNW metric:  
and BG metric:**

$$|q_{JNW}| \geq |\nu_{BG}| \quad \Rightarrow \quad |\Gamma_{JNW}| \geq |\Gamma_{BG}|$$

If the  $q$  parameter in the rotating JNW metric is greater than or equal to the  $\nu$  parameter in the BG metric, the BG metric is more stable compared to the rotating JNW metric.

## Charged three parameter metrics

**Einstein-Hilbert action in the presence of an electromagnetic field and a massless scalar field:**

$$S = \int d^4x \sqrt{-g} [R - F_{\mu\nu} F^{\mu\nu} + g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi].$$

**Field equations:**

$$R_{\mu\nu} = \partial_\mu \varphi(r) \partial_\nu \varphi(r) + T_{\mu\nu}^{EM},$$

$$\square \varphi(r) = 0,$$

$$T_{\mu\nu}^{EM} = 2 F_\mu{}^\alpha F_{\nu\alpha} - \frac{1}{2} g_{\mu\nu} F_{\gamma\sigma} F^{\gamma\sigma}.$$

# Charged three parameter metrics

**Ernst equations in the presence of the electromagnetic field:**

$$(Re(\varepsilon) + \Phi \Phi^*) \nabla^2 \varepsilon = \nabla \varepsilon \cdot (\nabla \varepsilon + 2 \Phi^* \nabla \Phi),$$

$$(Re(\varepsilon) + \Phi \Phi^*) \nabla^2 \Phi = \nabla \Phi \cdot (\nabla \varepsilon + 2 \Phi^* \nabla \Phi).$$

**Ernst potential:**

$$\varepsilon := f - |\Phi|^2 + i\chi$$

**$\Phi$  and  $\chi$  functions:**

$$\Phi := A_t + i \tilde{A}_\phi,$$

$$\hat{\phi} \times \nabla \tilde{A}_\phi := -\rho^{-1} f (\nabla A_\phi + \omega \nabla A_t).$$

$$\hat{\phi} \times \nabla \chi := -\rho^{-1} f^2 \nabla \omega - 2 \hat{\phi} \times \operatorname{Im}(\Phi^* \nabla \Phi).$$

## Charged three parameter metrics

LWP metric in  $(t, \rho, z, \phi)$  coordinate:

$$ds^2 = -f(dt - \omega d\phi)^2 + f^{-1} [e^{2\eta} (\rho^2 + dz^2) + \rho^2 d\phi^2].$$

three parameter metric in  $(t, r, \theta, \phi)$  coordinate:

$$ds^2 = -f^\gamma dt^2 + f^\mu k^\nu \left( \frac{dr^2}{f} + r^2 d\theta^2 \right) + f^{1-\gamma} r^2 d\phi^2.$$

# Charged three parameter metrics

**Harrison transformation:**

$$\varepsilon' = \frac{\varepsilon}{1 - 2\alpha^* \Phi - |\alpha|^2 \epsilon}, \quad \Phi' = \frac{\alpha \varepsilon + \Phi}{1 - 2\alpha^* \Phi - |\alpha|^2 \epsilon}.$$

**Metrics functions:**

$$\varepsilon = f(r) = \left(1 - \frac{2m}{r}\right)^\gamma,$$

$$\omega = \Phi = 0,$$

$$e^{2\eta} = \left[1 + \frac{m^2 \sin^2 \theta}{r(r-2m)}\right]^\nu \left[ \frac{2r(r-2m)}{m^2 [1 - \cos(2\theta)] + 2r[r-2M]} \right].$$

# Charged three parameter metrics

**Harrison transformation:**

$$\varepsilon' = \frac{1}{r^\gamma (r - 2m)^{-\gamma} - |\alpha|^2}, \quad \Phi' = \frac{\alpha}{r^\gamma (r - 2m)^{-\gamma} - |\alpha|^2}.$$

**Metrics functions:**

$$f' = \varepsilon' + |\Phi'|^2 = \frac{[r(r - 2m)]^\gamma}{[r^\gamma - (r - 2m)^\gamma - |\alpha|^2]^2},$$

$$\chi' = 0.$$

**Vector potential:**

$$A' = \frac{\alpha_R (r - 2m)^\gamma}{r^\gamma - |\alpha|^2 (r - 2m)^\gamma} dt + 2\gamma m \alpha_I \cos \theta d\phi.$$

## Charged three parameter metrics

### Transformations:

$$R = r(1 - |\alpha|^2) + 2m|\alpha|^2, \quad T = \left(\frac{t}{1 - |\alpha|^2}\right)^\gamma,$$
$$m = \frac{M}{1 + |\alpha|^2}, \quad q_m = 2m\alpha_I, \quad q_e = -2m\alpha_R.$$
$$e^2 := q_e^2 + q_m^2, \quad \xi := \frac{e}{M}, \quad p =: \sqrt{1 - \xi^2}.$$

# Charged three parameter metrics

**Final charged metrics:**

$$ds^2 = -\frac{\Delta^\gamma}{K^2} dt^2 + K^2 \Delta^\mu \Sigma^\nu \left[ \frac{1}{\Delta} dR^2 + R^2 d\theta^2 \right] \\ + R^2 K^2 \Delta^{1-\gamma} \sin^2 \theta d\phi^2,$$

**Metrics Functions:**

$$\Delta = 1 - \frac{2M}{R} + \frac{q_e^2 + q_m^2}{R^2},$$

$$\Sigma = 1 - \frac{2M}{R} + \frac{q_e^2 + q_m^2}{R^2} + \frac{(M^2 - (q_e^2 + q_m^2)) \sin^2 \theta}{R^2},$$

$$K = \left( \frac{1}{2p} \right)^\gamma (1+p)^{\gamma-1} \left[ (1+p) \left( 1 - \frac{M(1-p)}{R} \right)^\gamma \right. \\ \left. - (1-p) \left( 1 - \frac{M(1+p)}{R} \right)^\gamma \right].$$

# Charged three parameter metrics

**Vector potential:**

$$A' = \frac{F(R)}{K} dt + q_m \gamma \cos \theta d\phi,$$

$$F(R) = \frac{q_e}{(1+p)M} \left(1 - \frac{(1+p)M}{R}\right)^\gamma$$

**Scalar field:**

$$\square \varphi = 0, \quad \Rightarrow \quad \varphi(r) = \sqrt{\frac{1 - \gamma^2 - \nu}{2}} \ln \left( \frac{r - m - \sqrt{m^2 - q_e^2 - q_m^2}}{r - m + \sqrt{m^2 - q_e^2 - q_m^2}} \right).$$

# Charged three parameter metrics

**Ricci scalar:**

$$R(r, \theta) = \frac{Y(r, \theta)}{r^2 K^3 \Delta^{1+\mu} \Sigma^{2+\nu}}$$

**Singularities points:**

$$r_{singularity} = \begin{cases} 0 & \text{for all } \mu \text{ and } \nu \\ m \pm \sqrt{m^2 - (q_e^2 + q_m^2)} & \text{for } \mu > -1 \\ m \pm \sqrt{(m^2 - (q_e^2 + q_m^2)) \cos^2 \theta} & \text{for } \nu > -2 \end{cases}$$

## Charged three parameter metrics

**Charged metric in the presence of a scalar field at  $\gamma = 1$ :**

$$ds^2 = -\left(1 - \frac{2m}{r} + \frac{q^2}{r^2}\right) dt^2 + \left[1 + \frac{(m^2 - q^2) \sin^2 \theta}{r^2 - 2m + q^2}\right]^\nu dr^2 \\ \times \left[\frac{dr^2}{1 - 2m/r + q^2/r^2} + r^2 d\theta^2\right] + r^2 \sin^2 \theta d\phi^2.$$

In the absence of a scalar field ( $\nu = 0$ ), the above metric becomes the Reissner–Nordström metric.

## Conclusion:

- ✓ The JNW metric is not the most general solution for the Einstein-Hilbert action in the presence of a scalar field.  
We derived a more general three parameter solution.
  
- ✓ The Janis-Newman method is not a suitable method for obtaining the rotating JNW metric.  
We obtained a class of rotating metrics that include the rotating JNW metric using the Ernst's method.

- ✓ By using Ernst's equations in the presence of an electromagnetic field and also using Harrison's transformations, we were able to obtain the three-parameter metric charged form in the presence of a scalar field.
- ✓ A lot of new calculations can be done for the new three parameter rotating metrics in the future.

**Thank You!**